

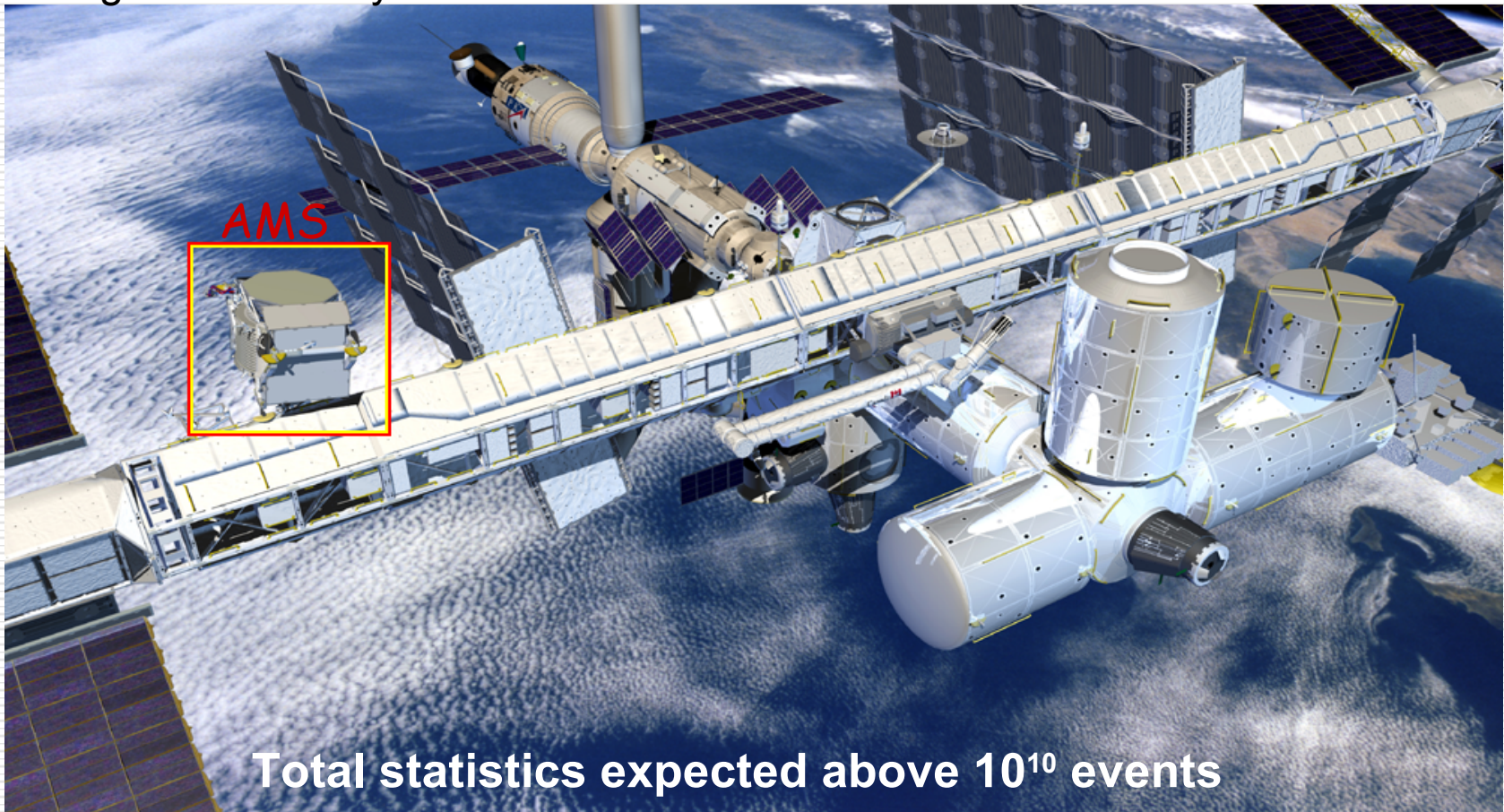
The Ring Imaging Cherenkov detector of the AMS experiment: test beam results with a prototype

Luísa Arruda

LIP - Lisbon

AMS-02 on the International Space Station

The **A**lpha **M**agnetic **S**pectrometer is a precision magnetic spectrometer scheduled to be installed in the **I**nternational **S**pace **S**tation (ISS) with a data taking of at least 3 years.



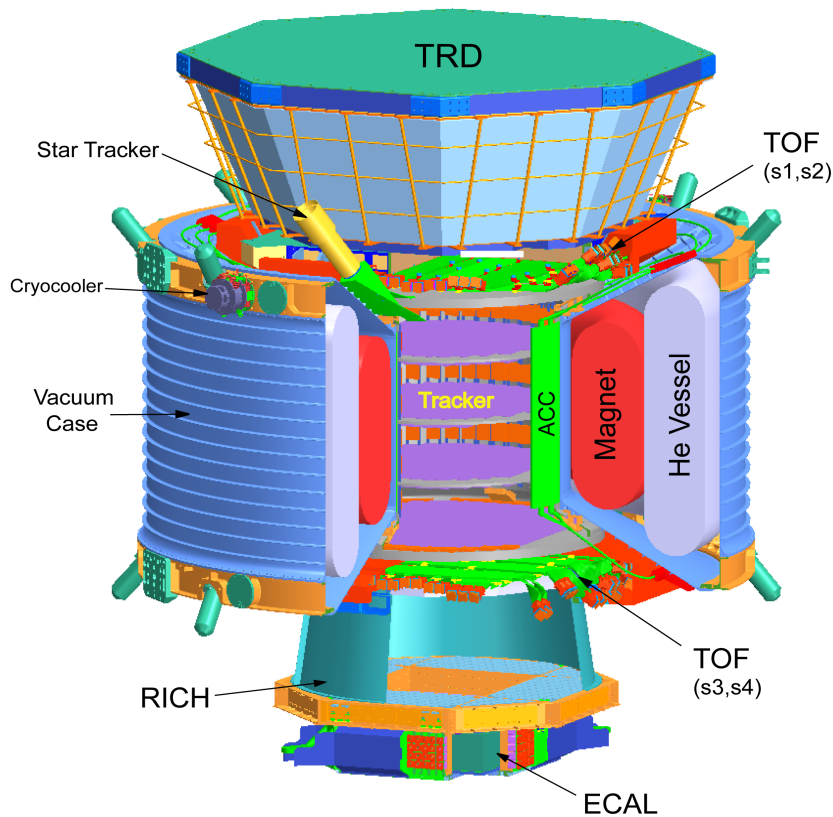
AMS-02 astrophysics aims

- **Search for cosmic antimatter**, through the detection of antinuclei with $|Z| \geq 2$; for antihelium nuclei the upper limit of detection will be $\overline{\text{He}}/\text{He} < 10^{-9}$;
- **Search for dark matter**. Neutralino annihilations may contribute with anomalies on different spectra: e^+ , \overline{p} , \overline{d} , γ
- **Precision measurements** on the relative abundance of different nuclei and isotopes of primary cosmic rays $E < 1$ TeV
 - ★ Secondary-to-primary ratios (B/C, ${}^3\text{He}/{}^4\text{He}$): test to propagation models
 - ★ Confinement times (${}^{10}\text{Be}/{}^9\text{Be}$): constraint to galactic halo models
 - ★ Long period of observation will give information on solar cycle variations

The AMS-02 experiment

AMS 02

(Alpha Magnetic Spectrometer)



Size: 3x3x3 m³

AMS on ISS for 3 years

Weight: ~7 Tons

R.Becker 09/05/03

AMS
Alpha
Magnetic
Spectrometer
Integration MIT

RB0305AMSdetector

➤ Particle bending

Superconducting magnet

➤ Rigidity (p/Z)

Silicon Tracker

➤ Particle direction

Time-Of-Flight, Tracker, RICH

➤ Velocity (β)

*RICH, Time-Of-Flight, Transition
Radiation Detector*

➤ Charge (Z)

RICH, Tracker, Time-Of-Flight

➤ Trigger

*Time-Of-Flight, ECAL, Anti-Coincidence
Counter*

Ring Imaging Cherenkov Detector (RICH)

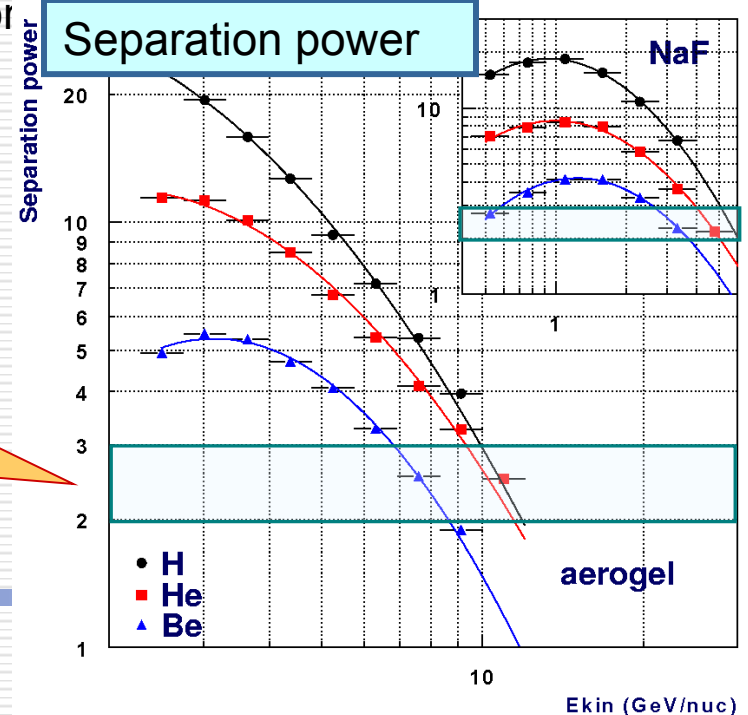
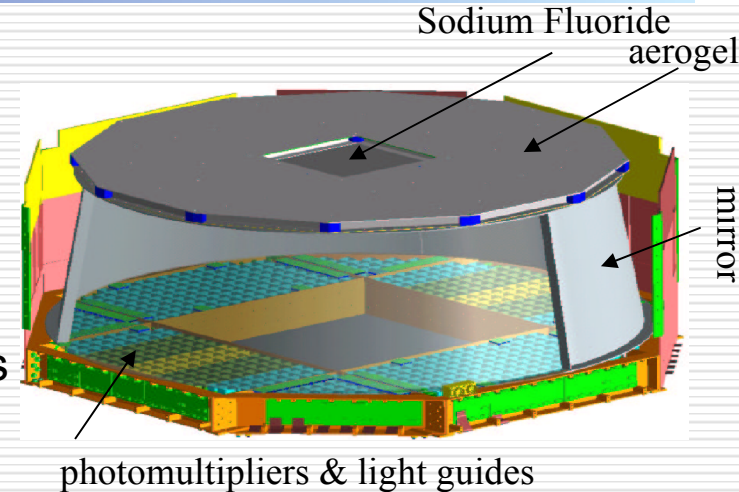
Physics Goals

- ✓ High accuracy on velocity measurement
 $\Delta\beta/\beta \sim 0.1\%$ for singly charged particles
- ✓ Electric charge measurement over a wide range of Z's
 At least up to iron element (Z=26)
- ✓ Contribution to AMS redundancy on albedo rejection

Simulated data samples: p, He, Be 

Separation criterium 

NaF can contribute at low energies but smaller acceptance!



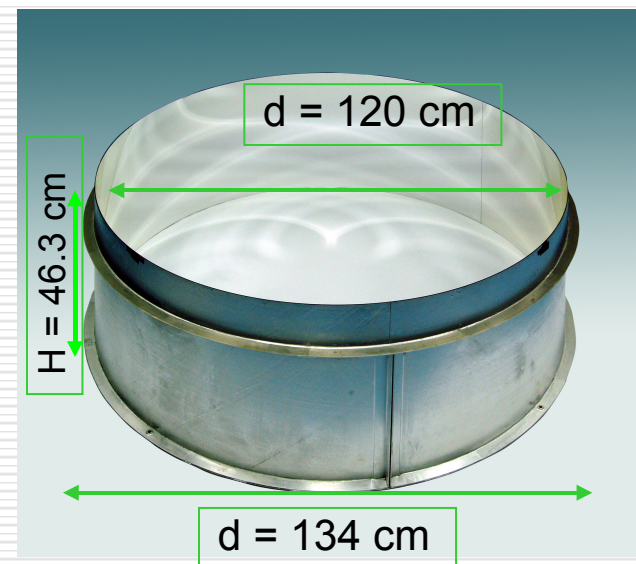
RICH detector: radiator plane and reflector

➤ Radiator:

➤ Dual solid radiator configuration

- increases the detector acceptance
- extends the kinematic coverage for lower energies

aerogel	1.050	2.5 cm	$p_{th} \sim 3 \text{ GeV}/c/\text{nuc}$
NaF	1.334	0.5 cm	$p_{th} \sim 1.1 \text{ GeV}/c/\text{nuc}$



➤ Conical reflector

- ~33% of the photons recovered
- carbon fiber structure with multilayer coating (Al+SiO₂)
- high reflectivity > 85% @ 420 nm

RICH detector: detection plane

➤ Photomultipliers

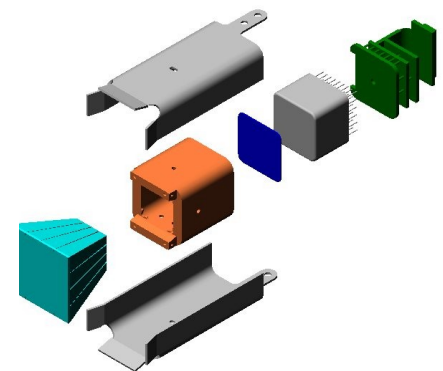
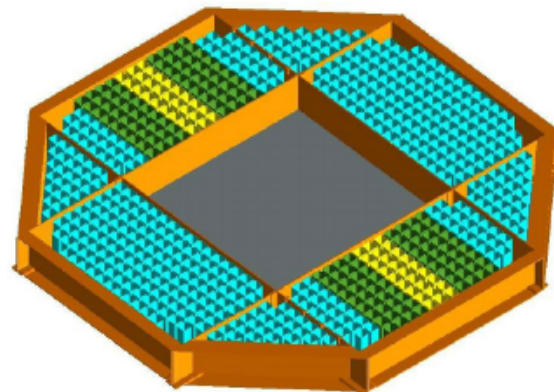
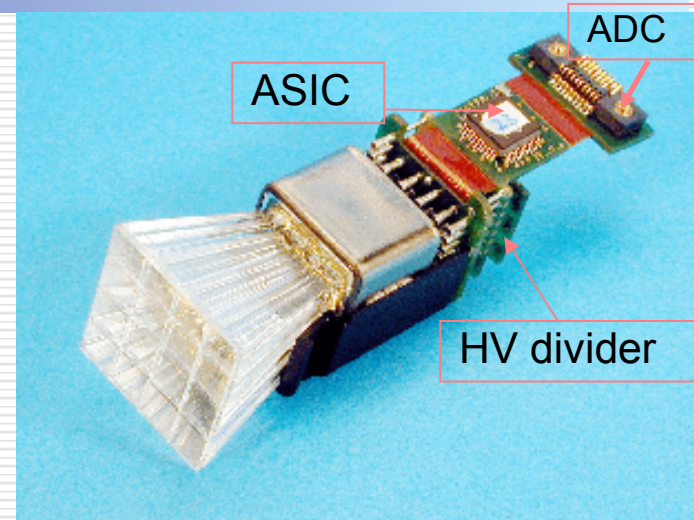
- matrix of 680 PMTs
- 4x4 multianode R7600-M16 (4.5 mm pitch)
- single photoelectron response
- spectral response 300-650 nm ($\lambda_{\max} \sim 420$ nm)
- high B_{stray} (~ 300 G) on readout plane
 - magnetic shielding needed (0.8-1.3 mm)

➤ Light Guides: increase photon collection eff

- Plexiglass (n=1.49) solid glass
- Spatial pixel granularity: 8.5x8.5 mm²

➤ Readout Electronics

- 16 channel ASIC developed
- two amplification gains (x 1,5)
- dynamic range from 1-100 pe



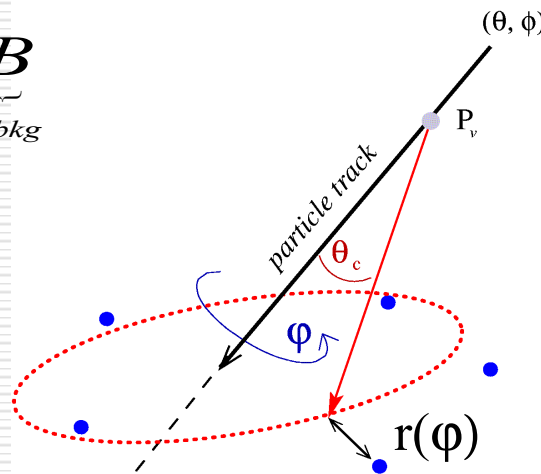
β and Z reconstruction algorithms

β reconstruction

✓ a method using all the hits with the maximization of a likelihood function providing the best θ_c angle

$$L(\theta) = \prod_{i=1}^{N_{hits}} P_i^{n_{pe_i}} \{r_i(\varphi_i(\theta))\}$$

$$P_i = \underbrace{\left(1 - \frac{b}{r_i}\right)}_{P_{signal}} S(r_i) + \underbrace{B}_{P_{bkg}}$$



Z reconstruction

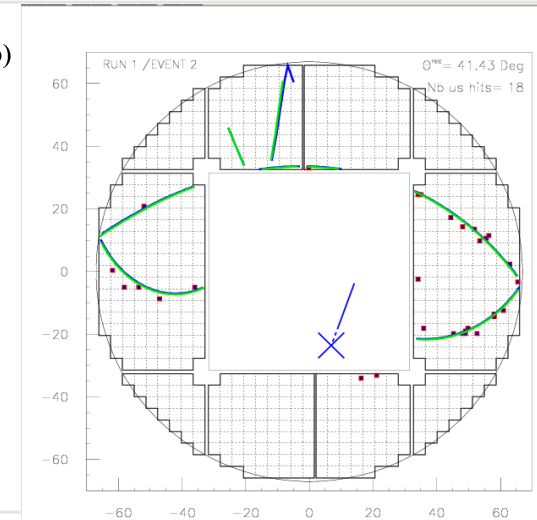
✓ A method based in

$$N_{pe} = N_0 Z^2$$



$$N_0 = N_{pe} (Z = 1)$$

N_0 semi-analytically evaluated



Test Beam 2003: experimental setup

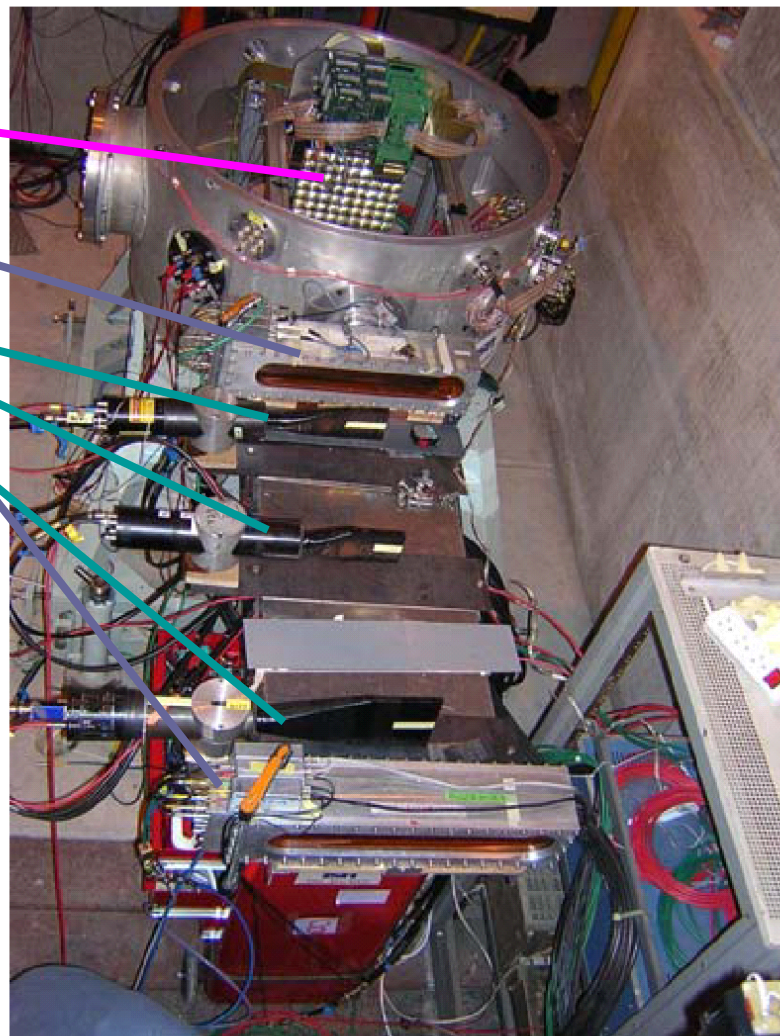
RICH prototype

MWPC

Organic scintillators &
Cherenkov counter

- Experimental area H8-SPS at CERN, October 2003
 - **Primary Beam:** *In*
 - **Target:** *Pb*
 - **Momentum:** *158 GeV/c/nucleon*
- Beam selection: *A/Z=2, 2.25, 2.35*

+ Tracker and TOF prototypes



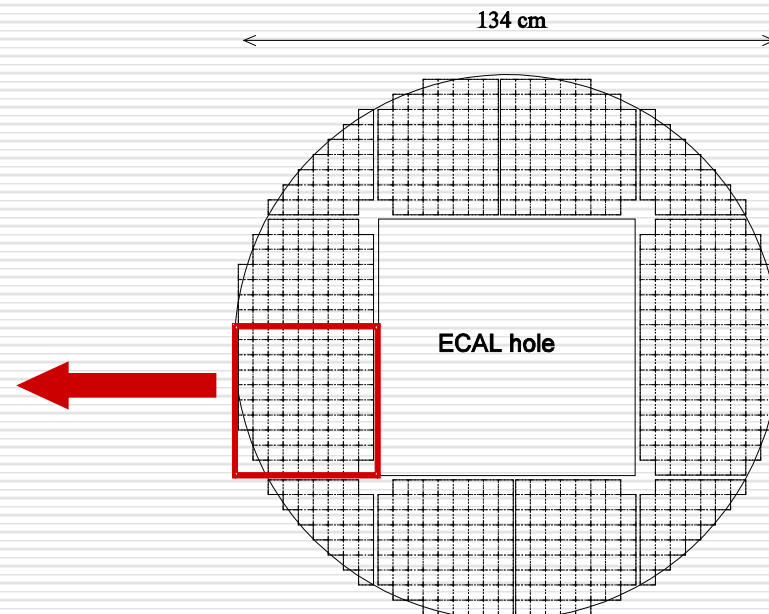
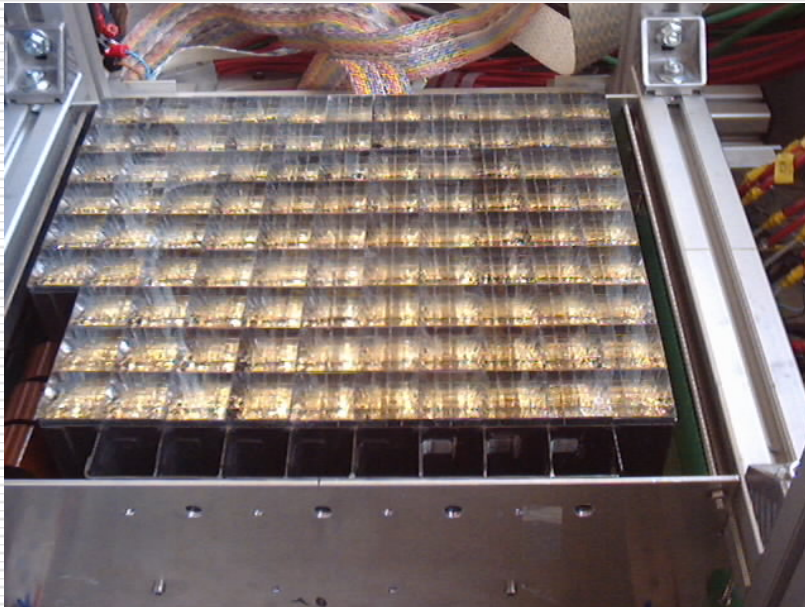
Test Beam 2003: experimental setup

RICH Prototype:

- detection matrix with 96 PMTs
- radiator: several tiles of aerogel, NaF
- mirror prototype with 1/12 of the final azimuthal coverage

Aims:

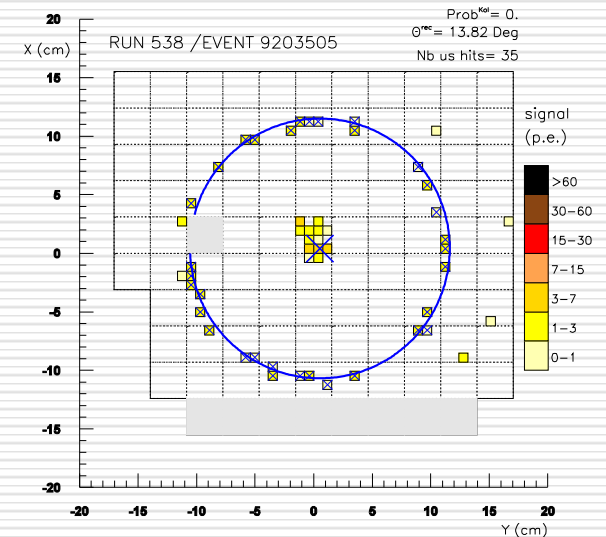
- ✓ testing flight front-end electronics;
- ✓ characterization of the AGL and NaF radiators;
- ✓ estimation of the mirror reflectivity;
- ✓ evaluation of the prototype global performance.



Test Beam 2003: analyzed data

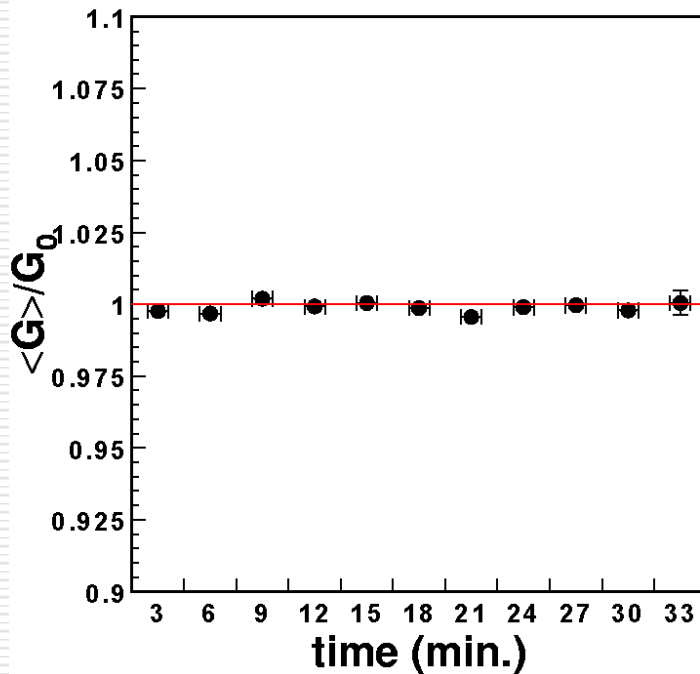
Radiators		
radiator	n	Thickness(cm)
Aerogel Novosibirsk (CIN103)	1.03	3
Aerogel Matsushita (MEC103)	1.03	3 X 1.1
Aerogel Novosibirsk (CIN105)	1.05	2.5
Sodium Fluoride	1.33	0.474

- 11 days of data taking
- $\sim 10^7$ events accumulated
- fully contained rings

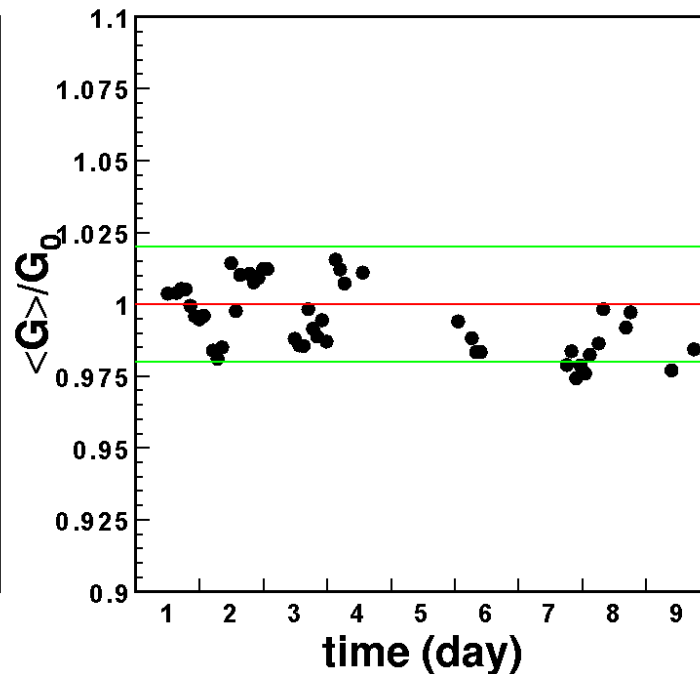


Detection matrix monitoring

- ✓ Photomultiplier gain stability monitored along the test beam.
- ✓ A gain stability within 2% was observed



At single run level



At test beam level

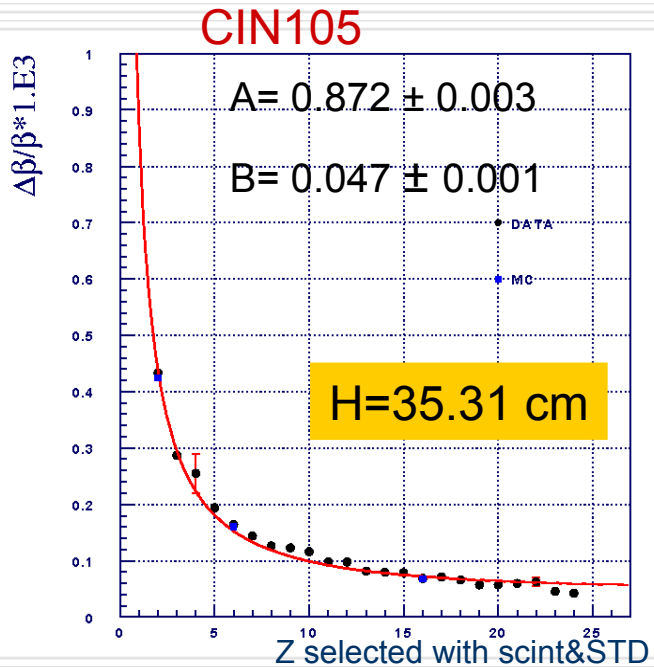
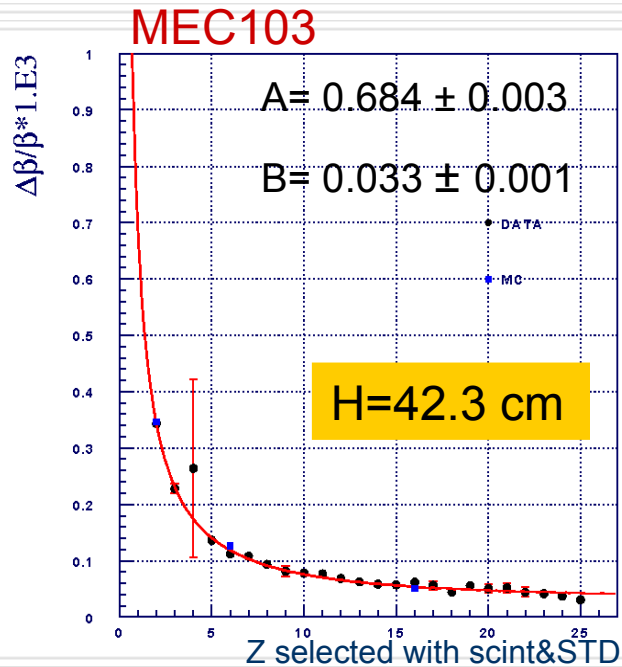
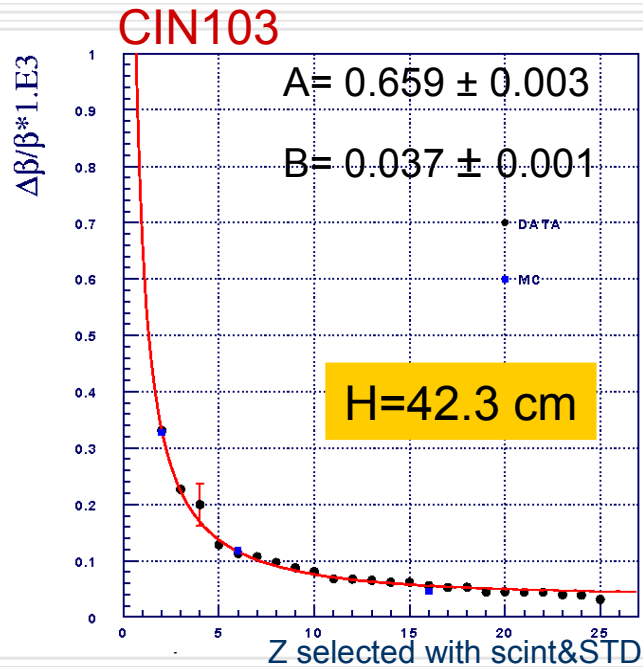
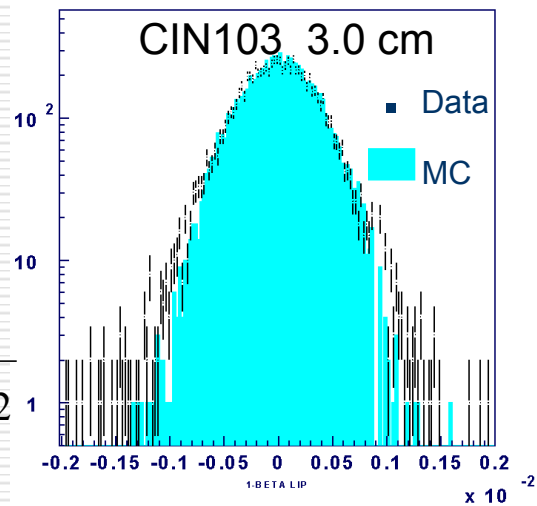
β resolution with the aerogel radiators

β resolution for Z=2, H=33.5 cm

radiator	CIN103	MEC103	CIN105
$\sigma(\beta) \times 10^3$	0.421 ± 0.003	0.435 ± 0.002	0.459 ± 0.004

All radiators fulfill the RICH requirements for β measurement !!

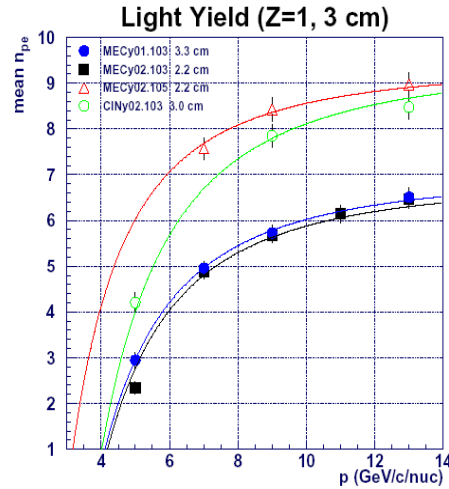
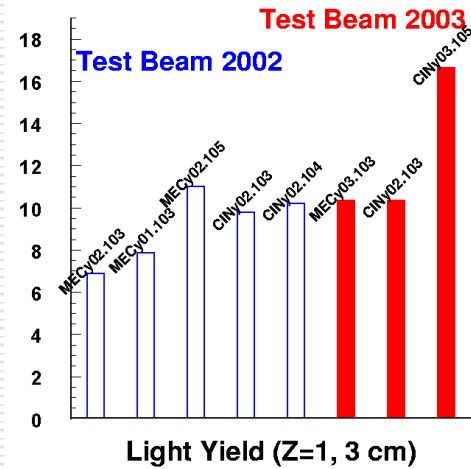
$$\frac{\Delta\beta}{\beta} = 10^{-3} \sqrt{\left(\frac{A}{Z}\right)^2 + B^2}$$



Photon yield and aerogel tile uniformity

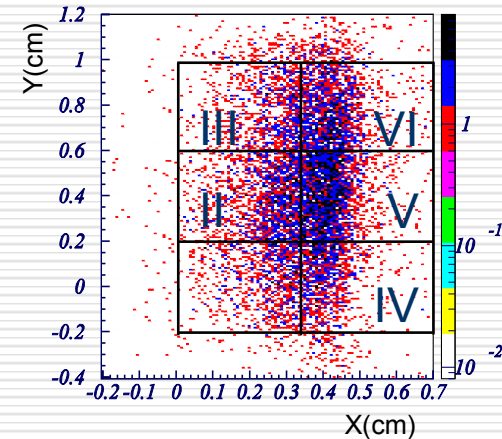
Light Yield

Tile Uniformity



Uniformity study based on data from a *wide beam*

Beam section $\sim 0.7 \times 1.2$ cm²



The aerogel of Novosibirsk 1.05 gives a significantly higher photon yield

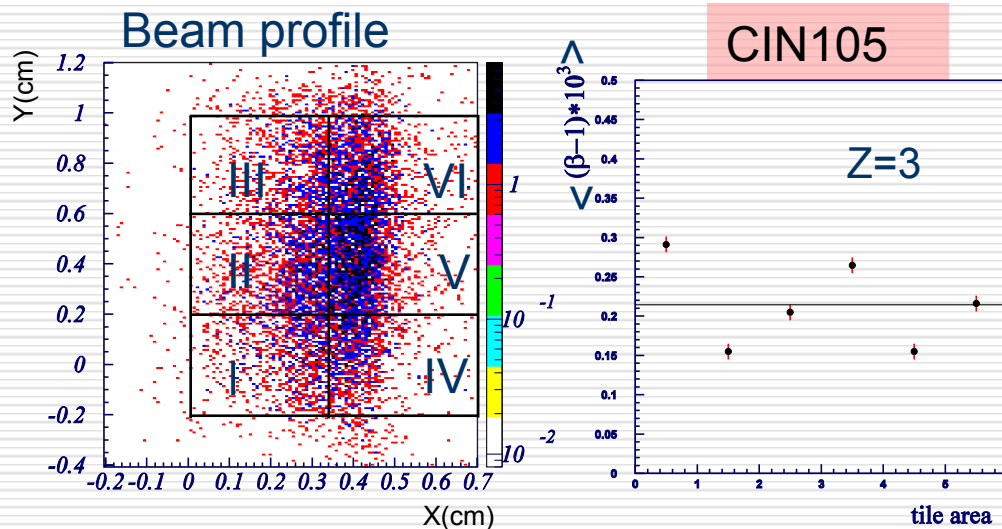
✓ photon yield

$$\Delta n_{pe} / n_{pe} < 1\%$$

Aerogel tile uniformity: refractive index

The study of bias in β mean value provides us with a direct estimation of the refractive index variation

$$\Delta\beta/\beta = \Delta n/n$$



Radiator	$\Delta n \times 10^3$ <i>Wide beam</i>
CIN105	0.06 ± 0.02

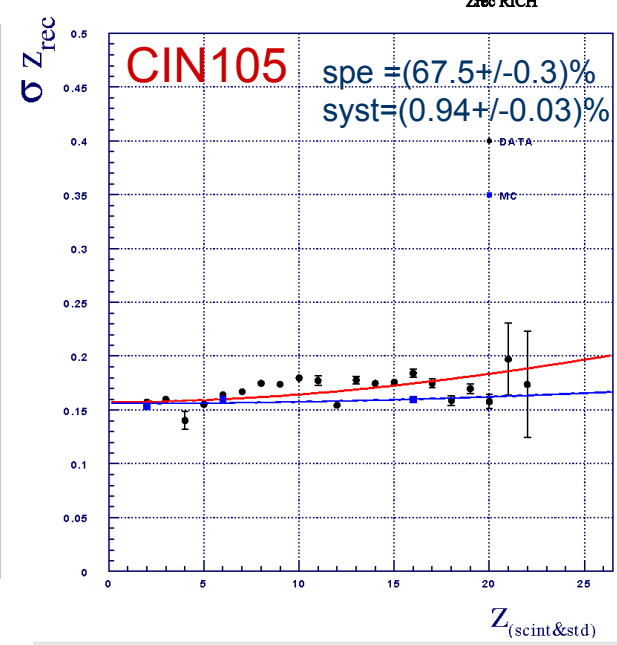
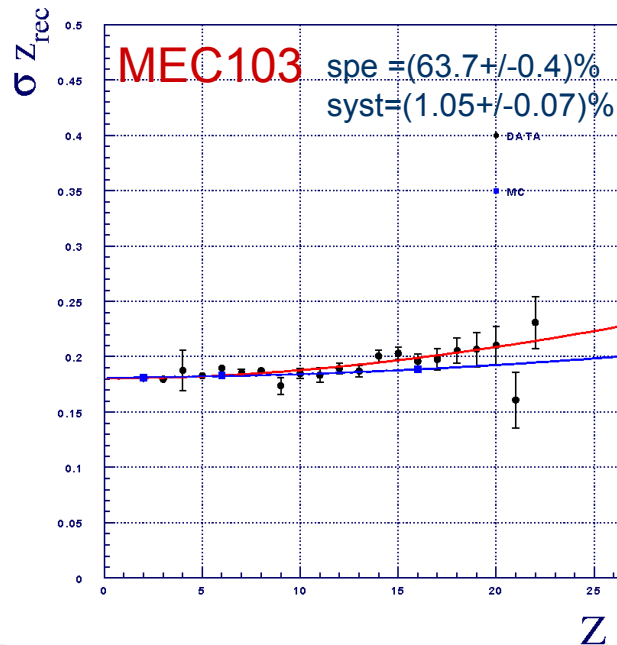
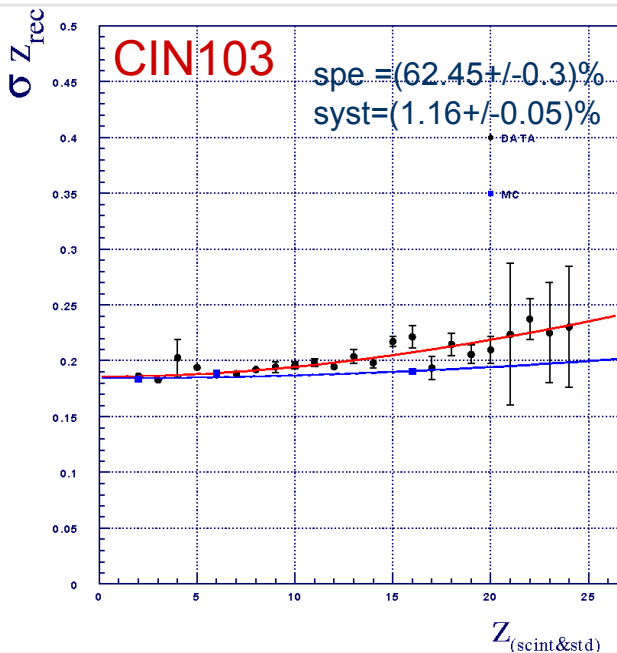
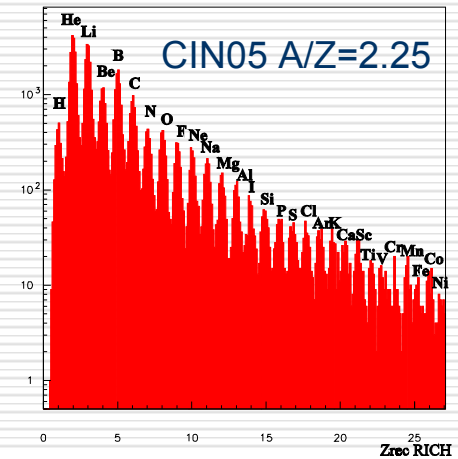
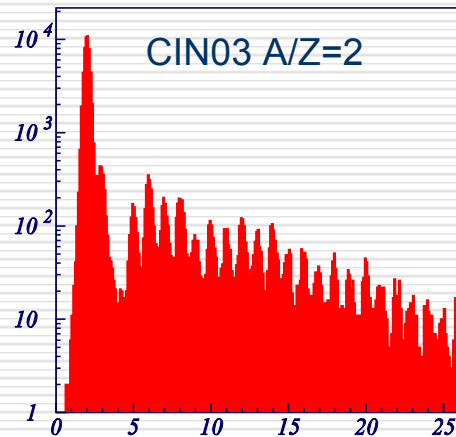
✓ refractive index
 $\Delta n < 10^{-4}$



CIN105 shows negligible non-uniformities in the refractive index

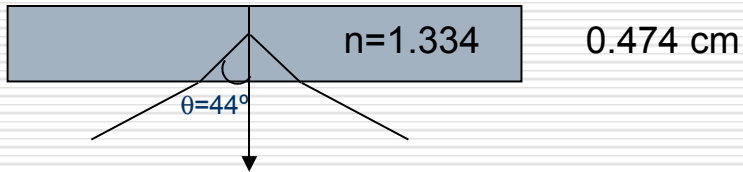
Z resolution with the aerogel radiators

$$\sigma(Z) = \frac{1}{2} \sqrt{\underbrace{\frac{1 + \sigma_{pe}^2}{N_0}}_{\text{stat}} + \underbrace{\left(\frac{\Delta N}{N}\right)_{\text{syst}}^2}_{\text{syst}}} Z^2$$



β reconstruction with the NaF radiator

NaF run 557 $\theta=0^\circ$

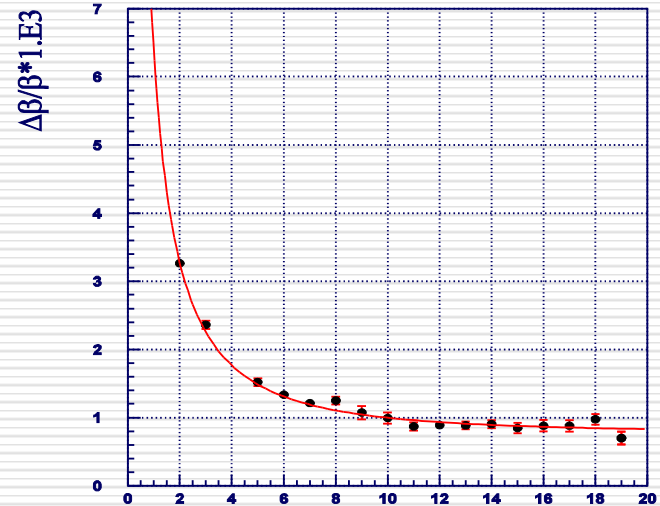
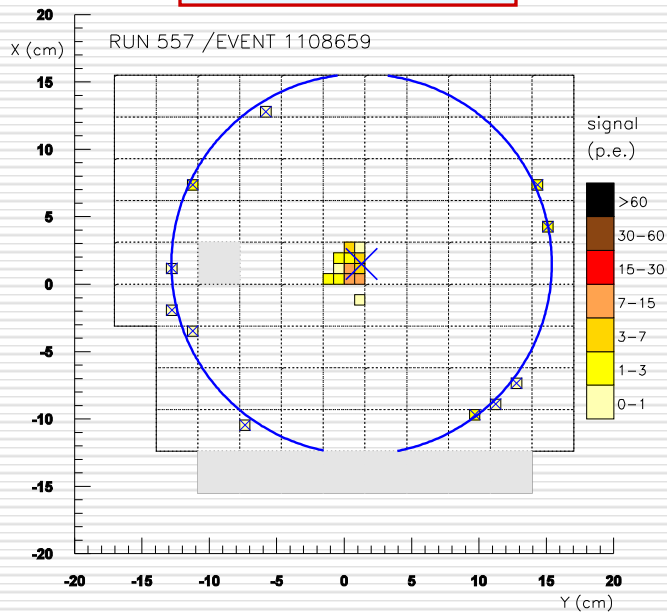


$$\frac{\Delta\beta}{\beta} = 10^{-3} \sqrt{\left(\frac{A}{Z}\right)^2 + B^2}$$

Helium event

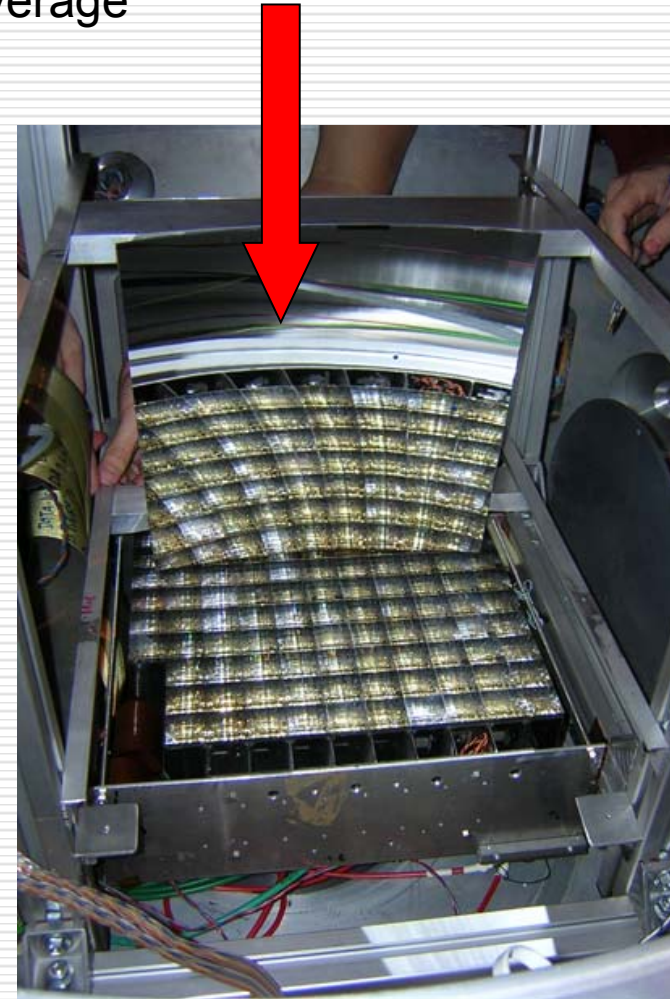
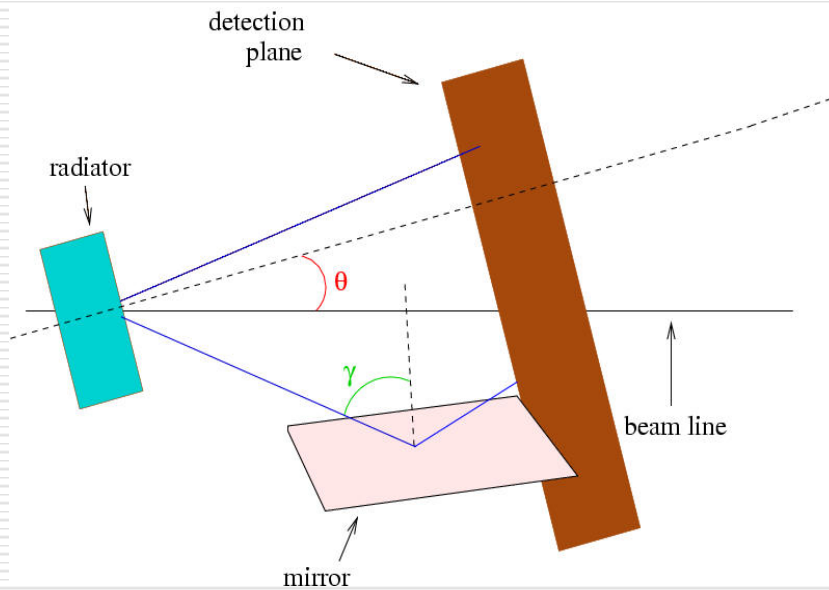
$$A=(6.19\pm 0.04)$$

$$B=(0.61\pm 0.02)$$

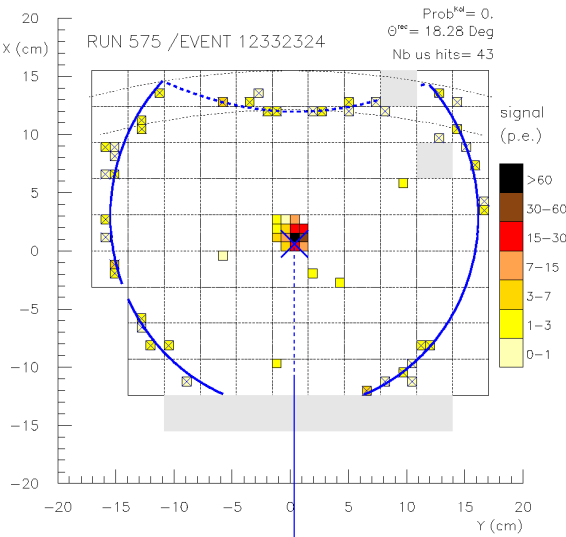


Mirror prototype

- Mirror segment with 1/12 of the total azimuthal coverage
- Data taken with different parameters:
 - refractive index (1.03, 1.05)
 - Different particle inclinations
 - Expansion height (42.3, 38.0 cm)



Mirror reflectivity



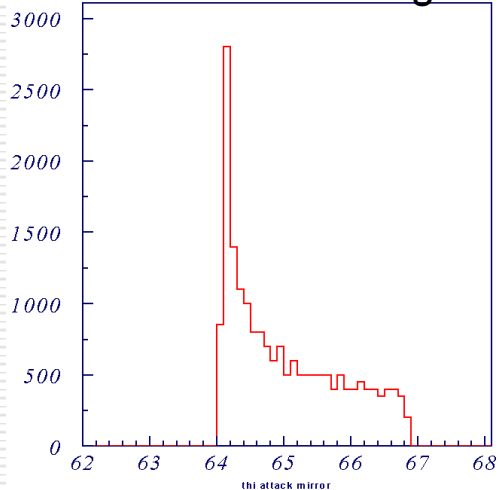
Helium events

$$\epsilon_{mir} = \frac{N_{pe}^{ref}}{N_{pe}^{dir}} \frac{\epsilon_{geo}^{dir}}{\epsilon_{geo}^{ref}} \frac{\epsilon_{lg}^{dir}}{\epsilon_{lg}^{ref}}$$

	Direct	Reflected
N_{pe}	35.9+/-0.1	9.51+/-0.02
ϵ_{LG}	0.7067+/-0.2E ⁻⁴	0.7709+/-0.3E ⁻⁴
ϵ_{geo}	0.6254+/-0.7E ⁻⁴	0.205+/-0.2E ⁻⁴

Reflectivity ~ (75.1 +/- 0.2) %

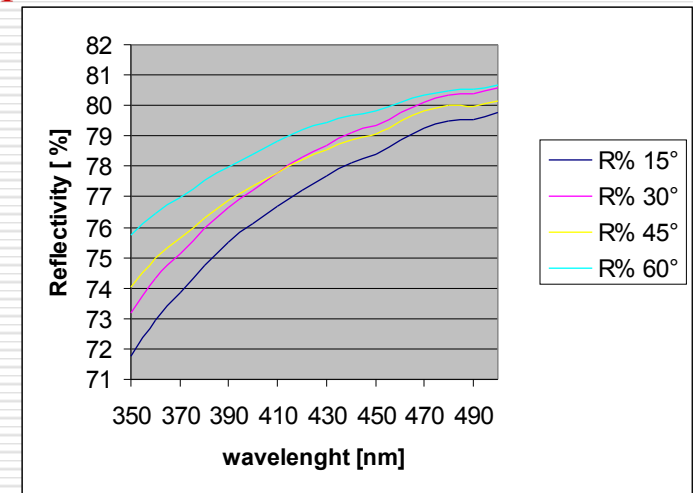
Photon incidence angle in the mirror



Mirror with a SiO coating not the final coating !!

Final sample reflectivity of ~85%

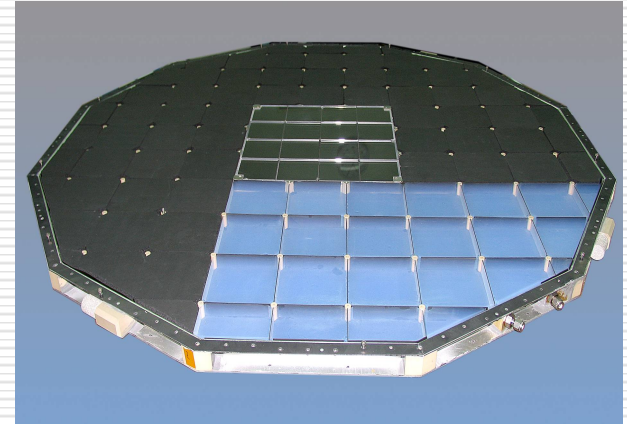
Optical measurement of the reflectivity



RICH status in October 2006

RICH Assembly has already started at CIEMAT, Madrid and is foreseen to be finished before July 2007

- ✓ Radiator:
 - ✓ aerogel tiles: optical characterization on going (@LPSC, Grenoble)
 - ✓ the aerogel container is ready (@CIEMAT, Madrid)
 - ✓ aerogel container + tiles: vibration starting today (@SERMS, Terni)
- ✓ Mirror: Prepared to be assembled. Already characterized with several measurements in LAB
- ✓ Detection plane:
 - ✓ 680 unit cells are finished and ready for grid assembly
 - ✓ 1st grid already assembled others are on going
 - ✓ functional tests of the unit-cells and full characterization done



Conclusions

- ✓ The AMS-02 experiment to be installed in the International Space Station will be equipped with a RICH which is foreseen to be completely assembled by July 2007
- ✓ The design of the RICH detector was validated through intensive tests of a prototype.
- ✓ Velocity and charge reconstruction algorithms were tested.
- ✓ Full agreement between the MC and data taken with the aerogel radiator.
- ✓ Mirror sample used in the test:
 - ✓ Reflectivity measurement with data in agreement with the measurements of the manufacturer.