

# **First look at the RICH Prototype test beam data**

**C. Delgado, E. Lanciotti, C. Palomares, F. Giovacchini,  
N. Sevilla, A. Torrentó, E. Sánchez, J. Casaus**

**RICH MEETING, CERN, 04/12/2002**

# OUTLINE

- Setup description
- Data sample
- Calibration procedure
- Event selection
- Preliminary results
- Next steps
- Conclusions



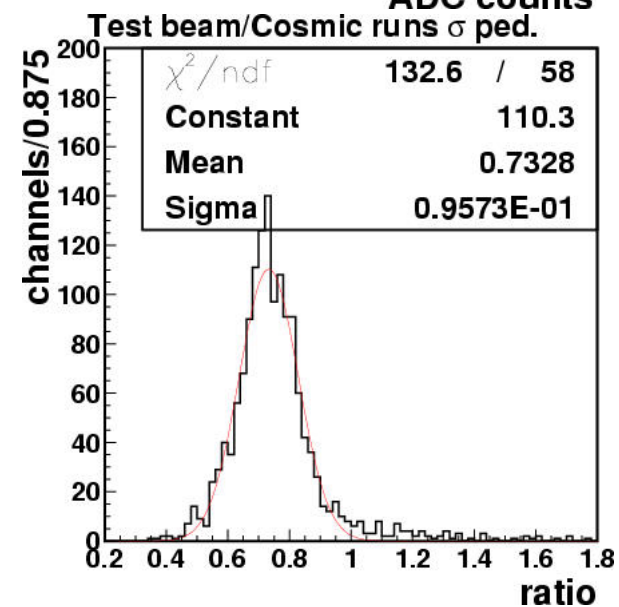
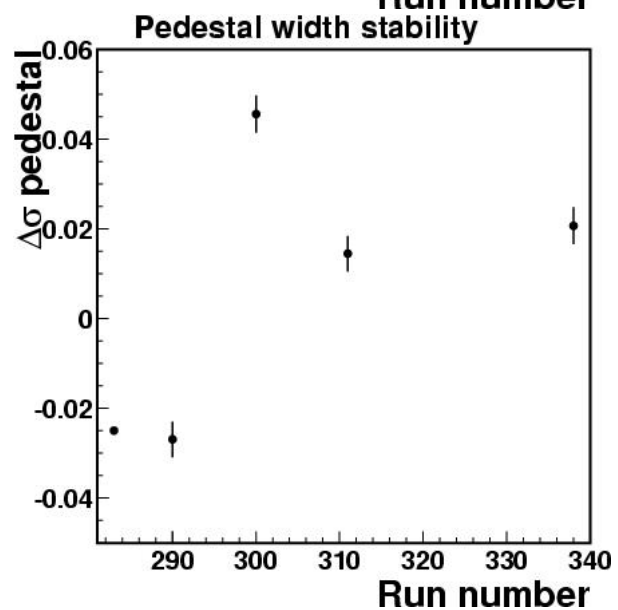
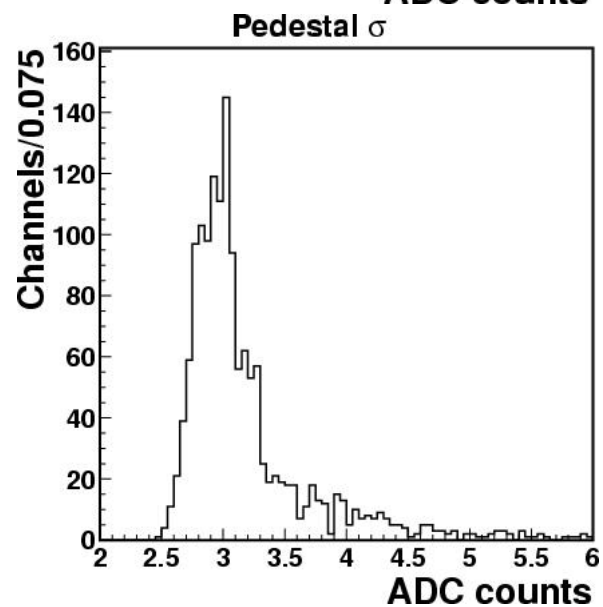
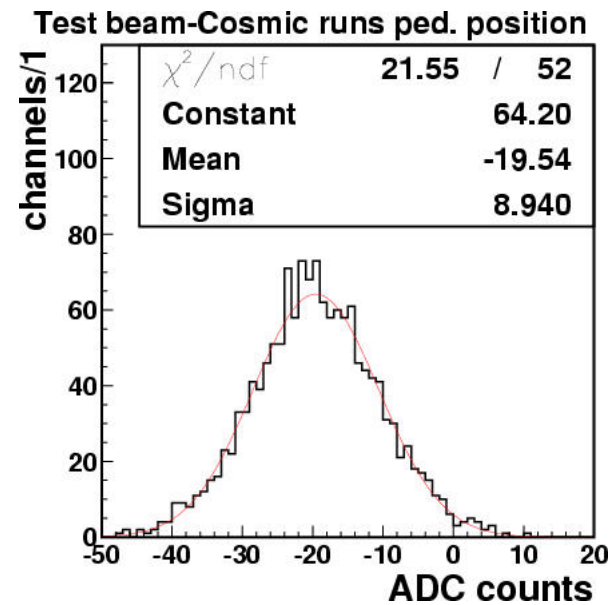
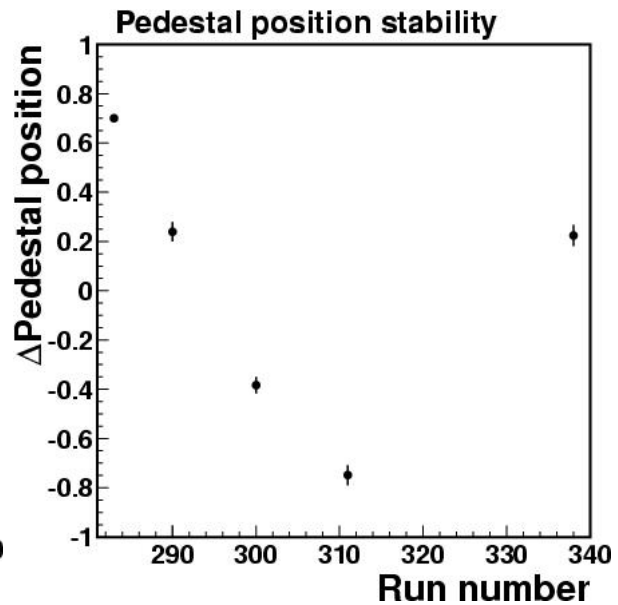
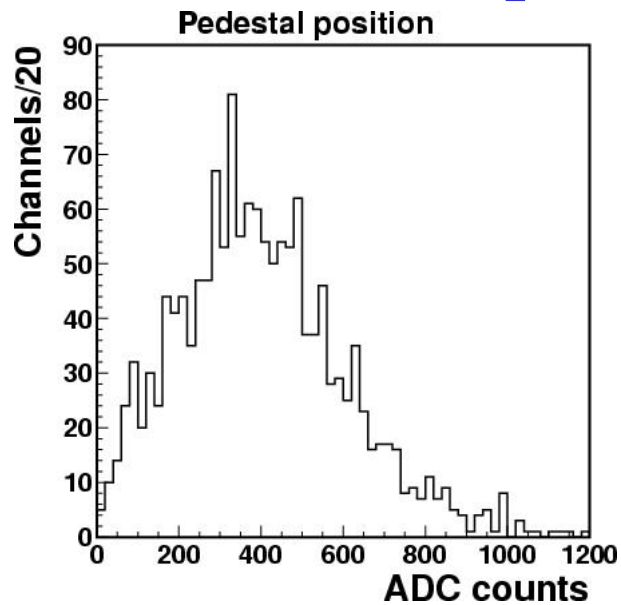
# DATA SAMPLE

- Data sample taken 18/10-19/10 (WC quality, event shift...)
- Proton,  $^3\text{He}$ ,  $A/Z=2$  runs selected
- Maximum 50k events/run processed

Run	Day	Init	End	Drift	B\rho	#Events	Beam Type
<b>New Matsushita</b>							
		1.03	2x10 mm				
RUN	DAY	INIT	END	DRIFT	B\rho	#EVENTS	BEAM TYPE
292	18/10	06:58	08:00	417.4	40	171135	A/Z=2
295	18/10	11:45		417.4	11	52000	p
<b>New Matsushita</b>							
		1.05	2x10 mm				
RUN	DAY	INIT	END	DRIFT	B\rho	#EVENTS	BEAM TYPE
302	18/10	15:35		375	13	39961	p
304	18/10	16:15		375	9	16390	A/Z=2
308	18/10	19:24	19:53	375	40	85023	A/Z=2 / Q15=0.2
<b>Novosibirsk-hydrophil.</b>							
		1.03	30 mm				
RUN	DAY	INIT	END	DRIFT	B\rho	#EVENTS	BEAM TYPE
310	18/10	21:24	21:40	417.4	40	74252	A/Z=2
312	18/10	21:54	21:59	417.4	30	71690	A/Z=3/2 He3
313	18/10	22:09	22:26	417.4	13	33688	p
<b>Novosibirsk-hydrophob.</b>							
		1.03	30 mm				
RUN	DAY	INIT	END	DRIFT	B\rho	#EVENTS	BEAM TYPE
285	17/10	23:00		417.4	40	117737	A/Z=2
325	19/10	07:16	07:25	417.4	13	30546	p
326	19/10	07:33	07:43	417.4	30	132132	A/Z=3/2 He3
327	19/10	07:54	08:	417.4	40	118633	A/Z=2
<b>Old Matsushita</b>							
		1.03	3x10 mm				
RUN	DAY	INIT	END	DRIFT	B\rho	#EVENTS	BEAM TYPE
319	19/10	02:55	03:13	417.4	13	35310	p
320	19/10	02:32	03:36	417.4	30	130557	A/Z=3/2 He3
321	19/10	03:46	04:25	417.4	40	135225	A/Z=2

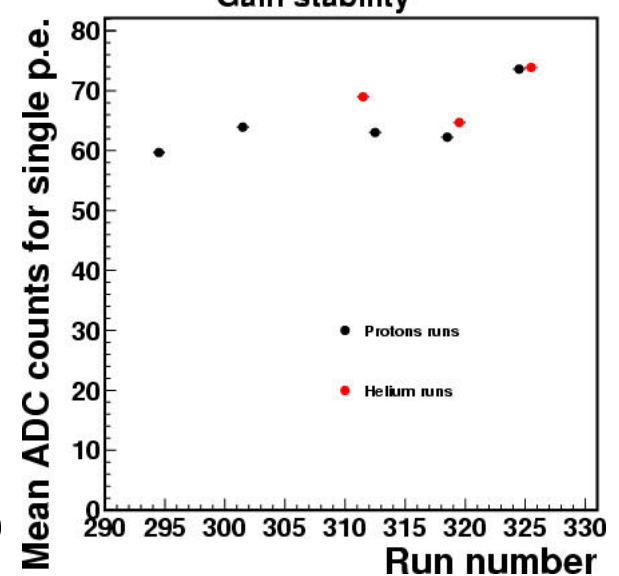
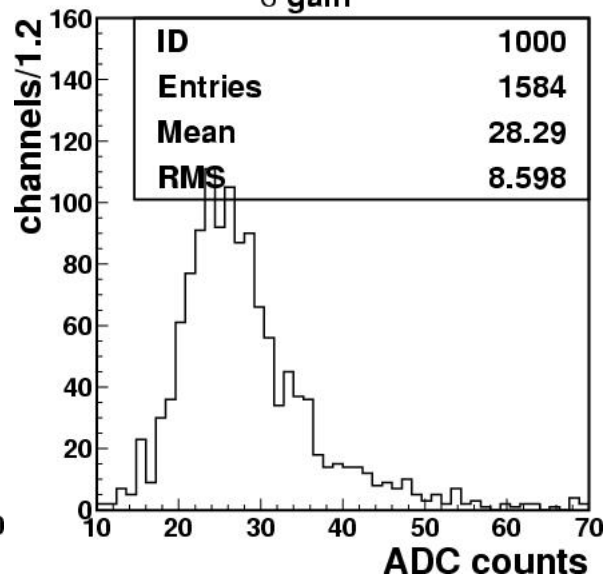
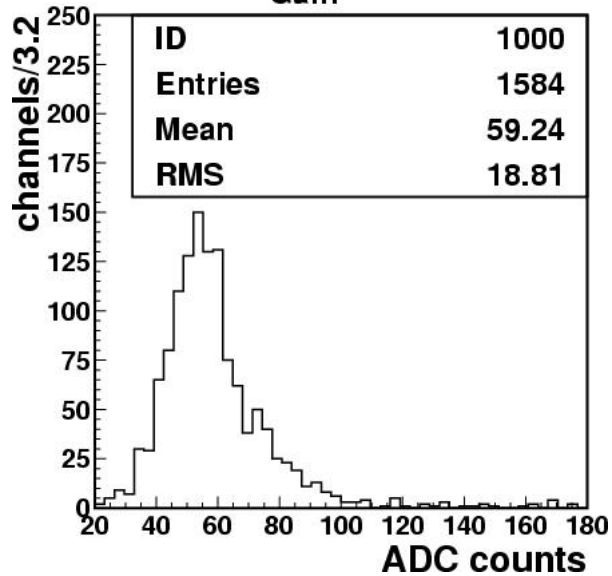
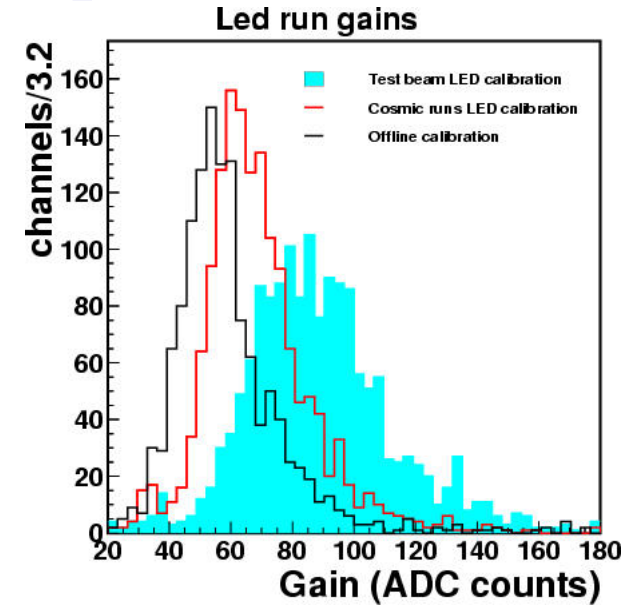
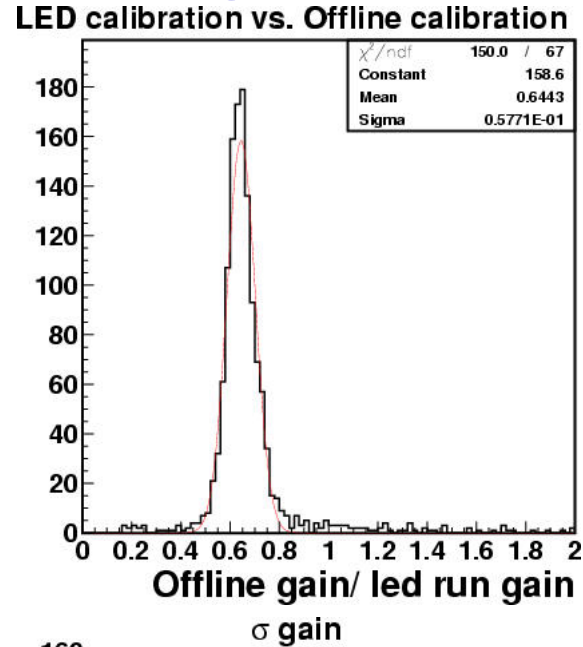
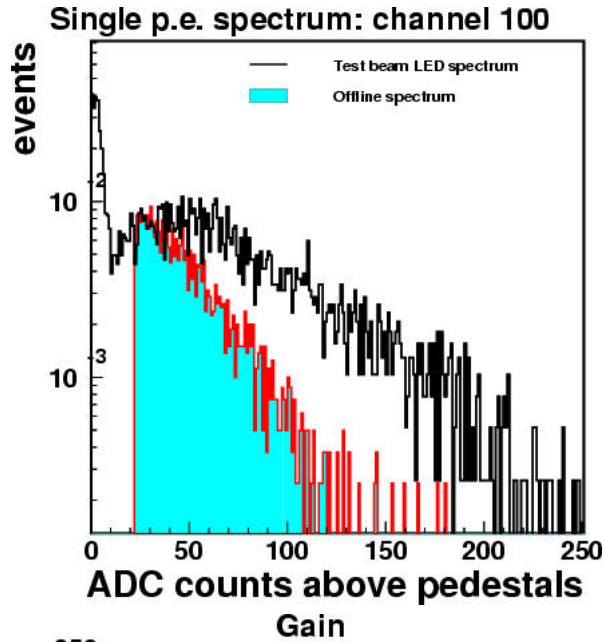
# PMT CALIBRATION

## ➤ Pedestals: 5 pedestal runs along the selected period



# PMT CALIBRATION

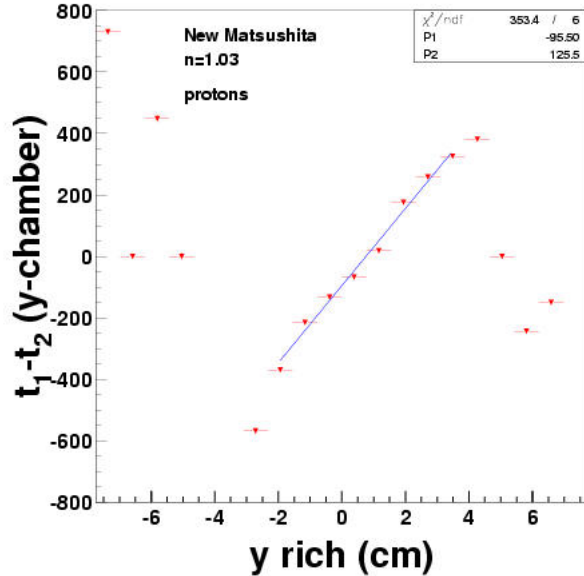
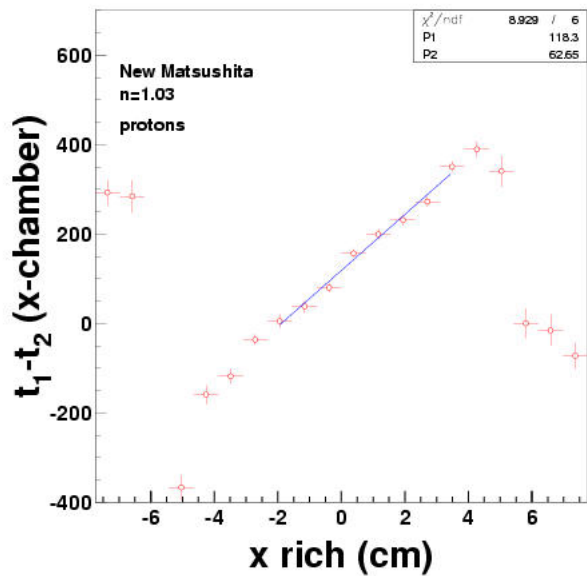
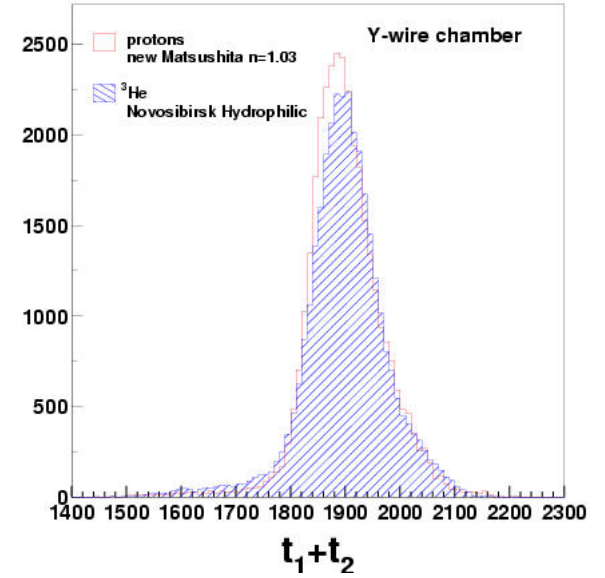
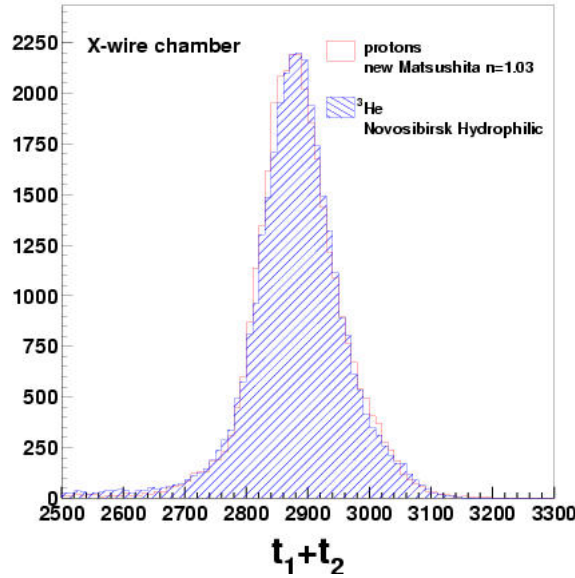
## ➤ Gain: 5 data runs along the selected period



# WC CALIBRATION

$$2500 < t_1^x + t_2^x < 3250$$

$$1400 < t_1^y + t_2^y < 2400$$



$$t_1^x - t_2^x = 116 + 64 x_{\text{RICH}}$$

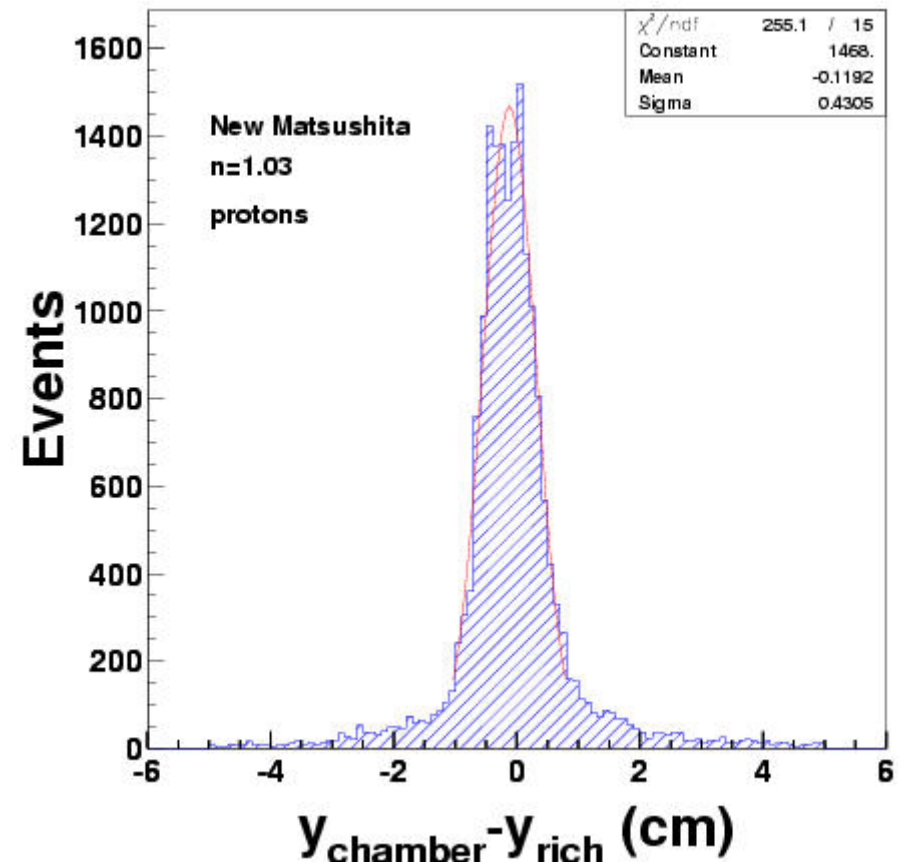
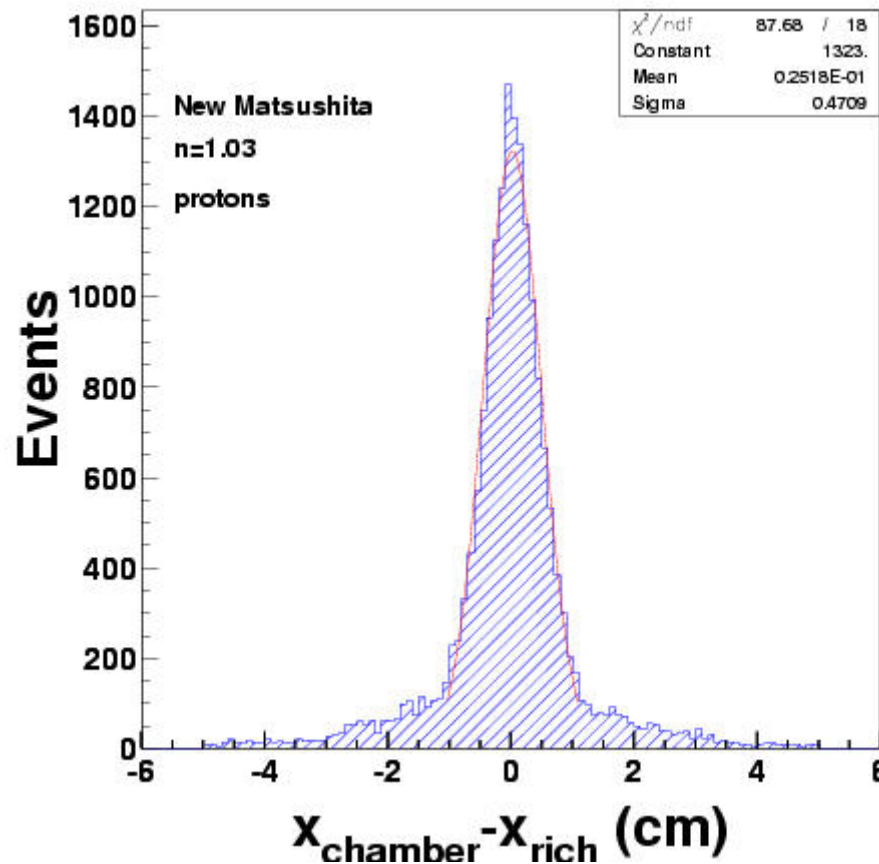
$$t_1^y - t_2^y = -95 + 127 y_{\text{RICH}}$$

# WC CALIBRATION

WC – RICH position residues (4.5 mm):

1. same in x & y axes
2. similar to those obtained @ ISN with cosmics (6mm)

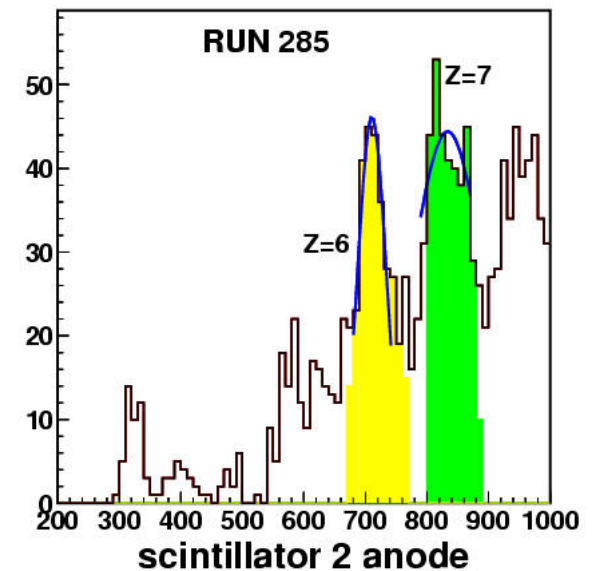
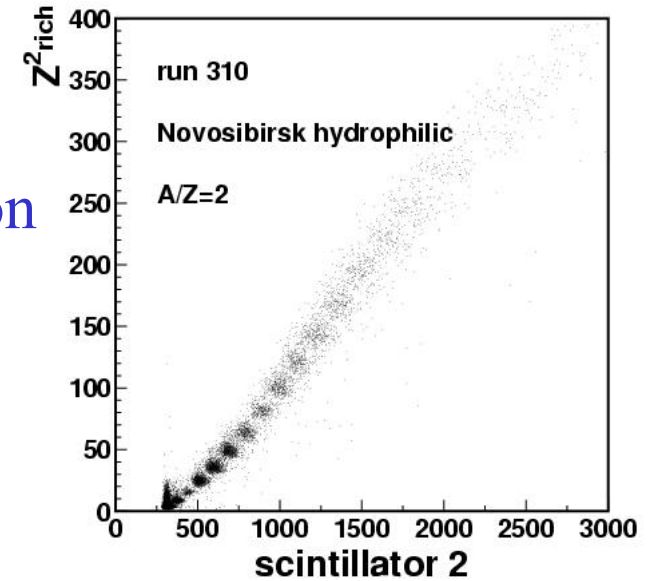
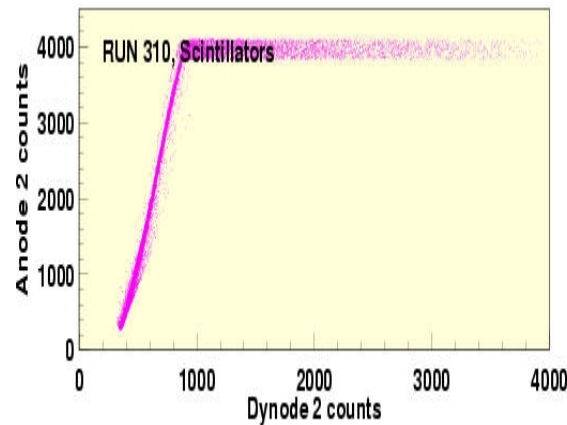
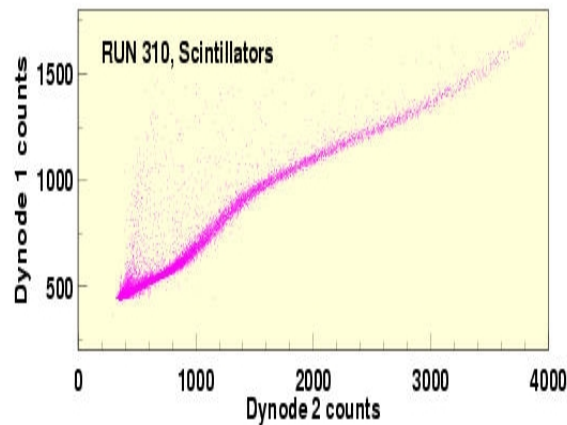
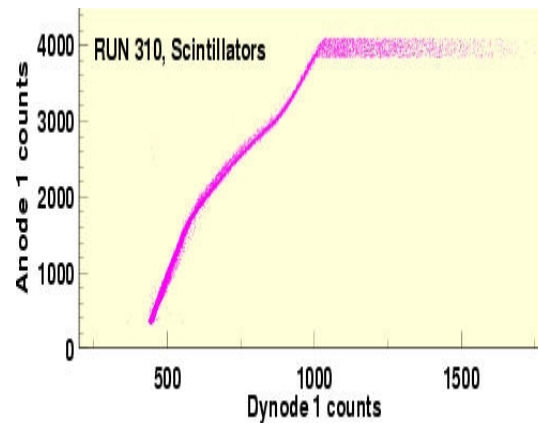
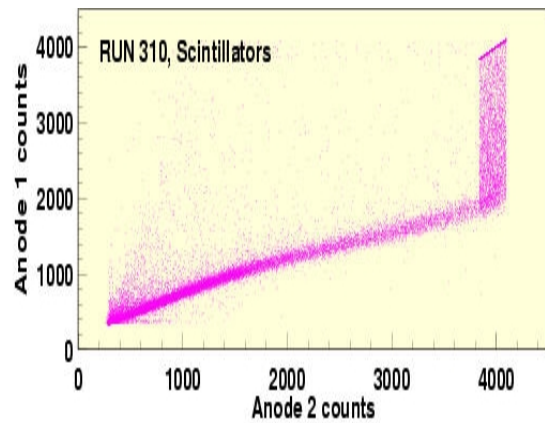
(RICH pixel contribution  $\approx 3.5$  mm)





# SCINT CALIBRATION

- Good correlation observed btwn SC1 & SC2
- Non linearities in anode & dynode signals
- Scint response calibrated using RICH selection

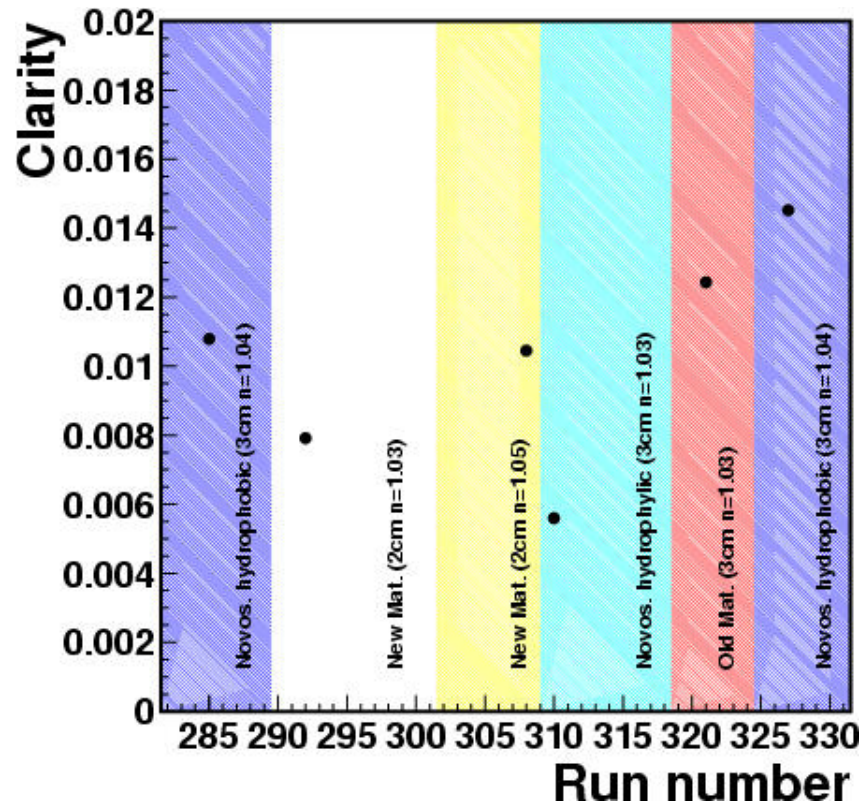
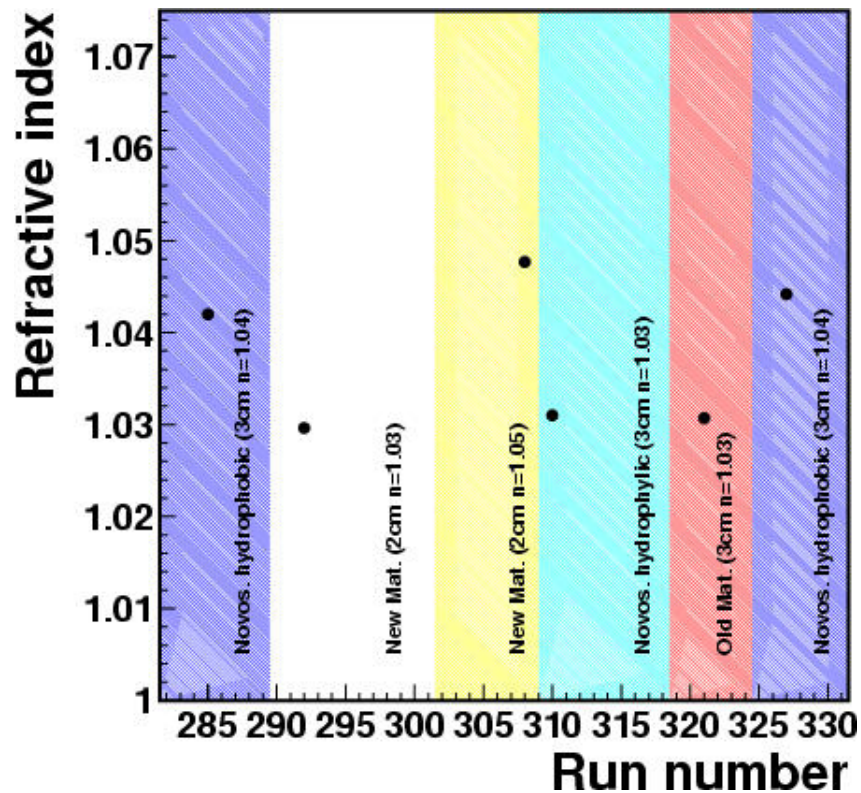


# RADIATOR PARAMETERS

Optical parameters can be (ambiguously!) tuned from the RICH velocity and charge reconstruction:

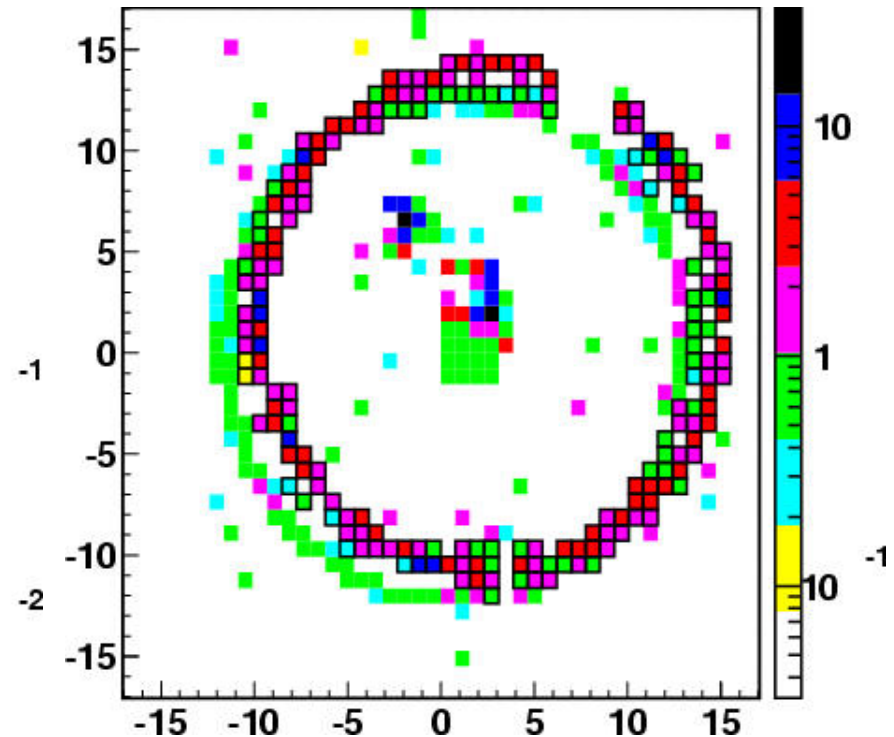
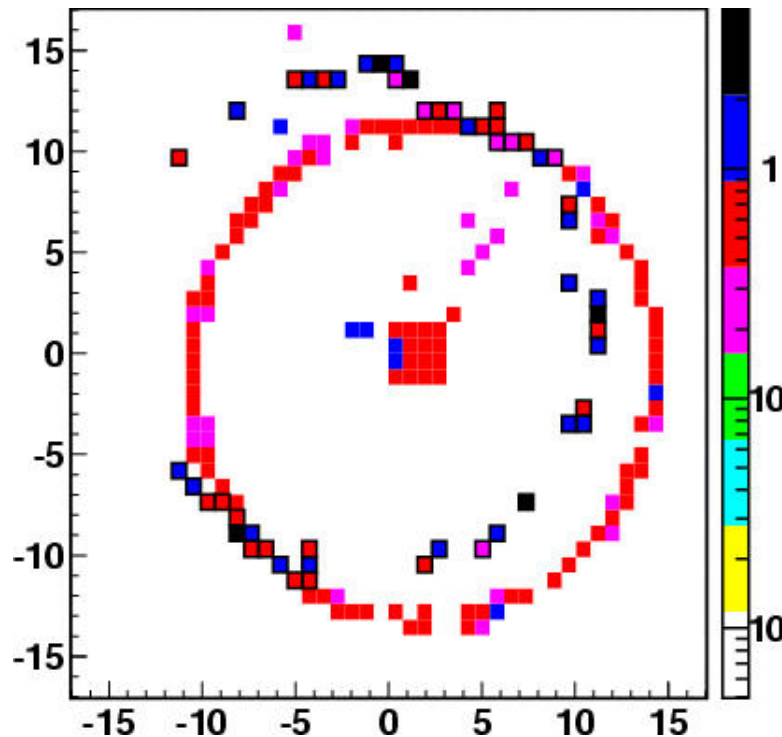
$\beta \Rightarrow$  Refraction index (or expansion length)

$Z \Rightarrow$  Clarity (or absorption length)



# EVENT SELECTION

- Reconstructed WC-track
- Loose WC-beam and WC-RICH matching
- $N_{\text{USED}} > 2$ ,  $N_{\text{EXP}} > 2$
- Multi-particle event rejection
  - Multi-Ring (Charge/ChargeInRing < 1.75)
  - Multi-Track (ChargeSingleHit < 2.3 x ChargeInRing + 2.7)



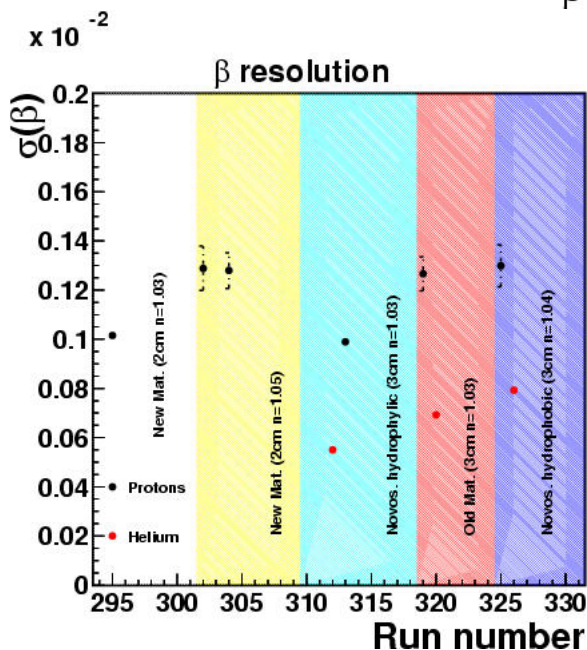
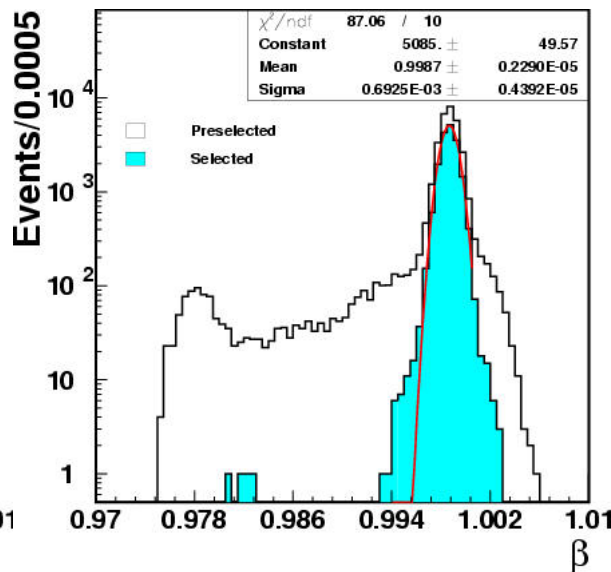
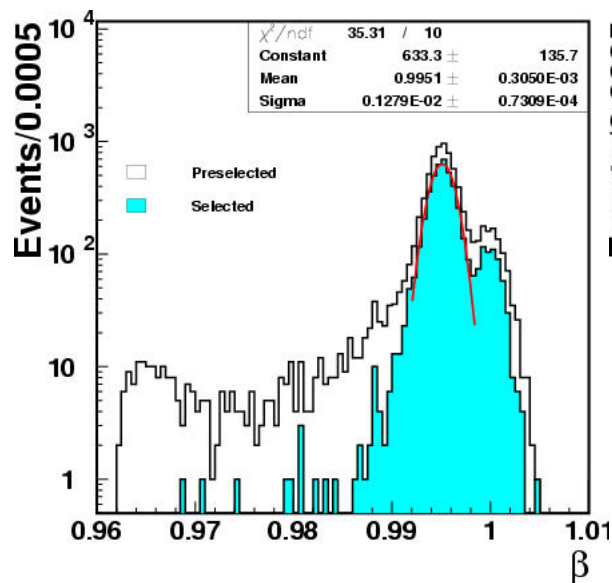
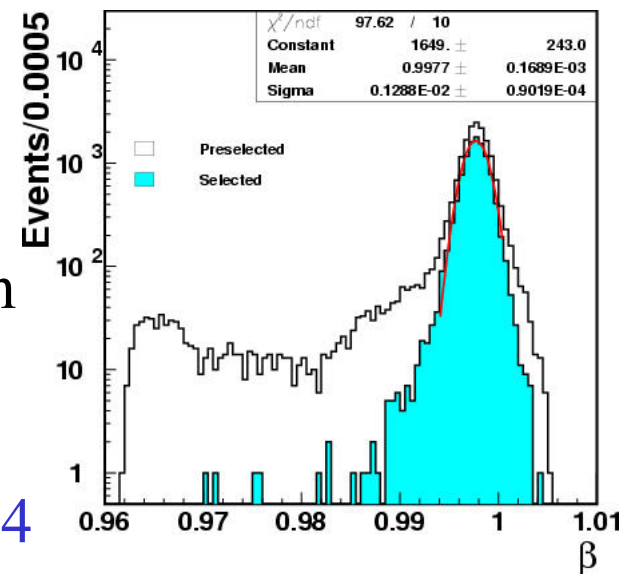
# PRELIMINARY RESULTS

## $\beta$ determination

After selection, tails are mainly coming from multi-particle events and beam contamination

$$\sigma(\beta)/\beta \approx 1.0\% : \text{NM } 1.03, \text{NV } 1.03$$

$$\sigma(\beta)/\beta \approx 1.3\% : \text{NM } 1.05, \text{OM } 1.03, \text{NV } 1.04$$



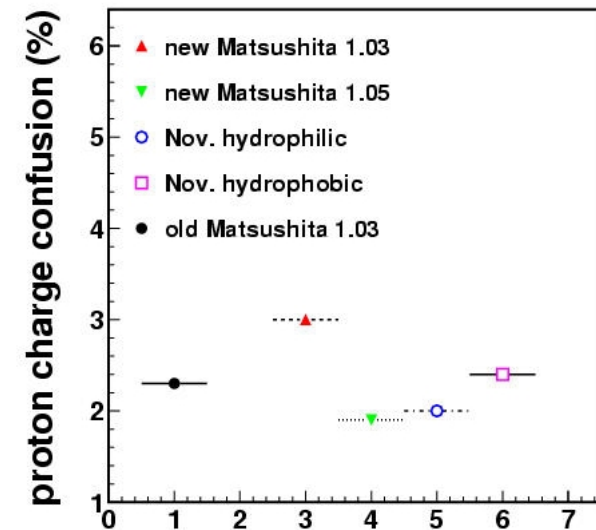
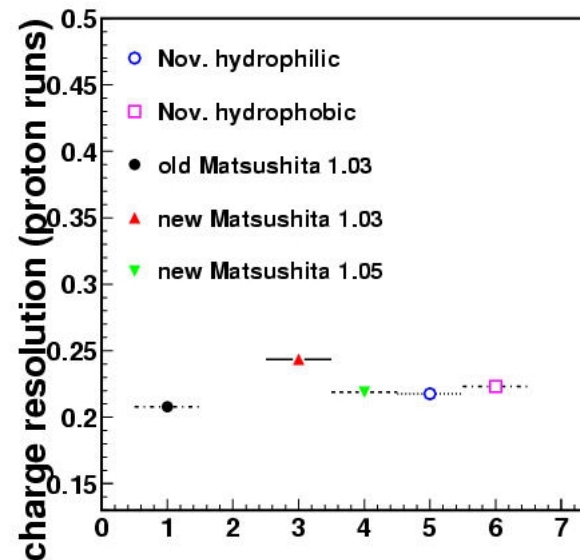
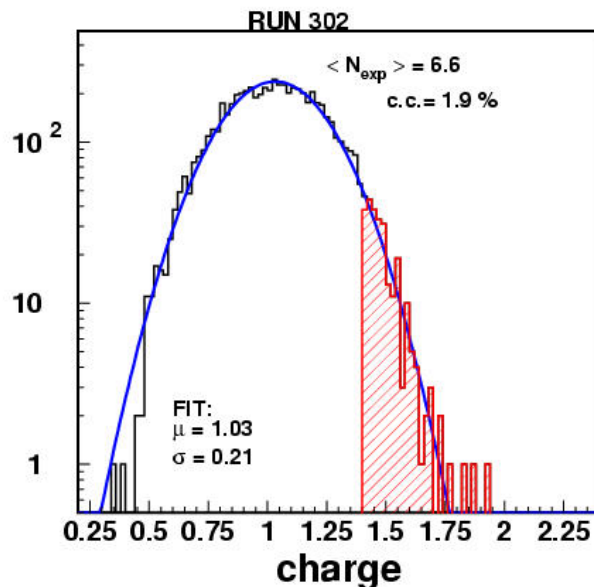
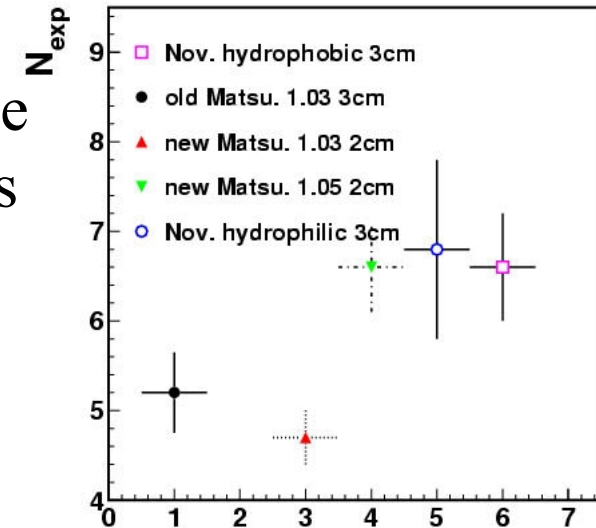
# PRELIMINARY RESULTS

## Z determination

Since the rings are completely contained in the PMT matrix, the optimum charge resolution is obtained for every radiator.

$$N_{\text{EXP}} \approx 7 : \text{NM } 1.05, \text{NV } 1.03, \text{NV } 1.04$$

$$N_{\text{EXP}} \approx 5 : \text{NM } 1.03, \text{OM } 1.03$$

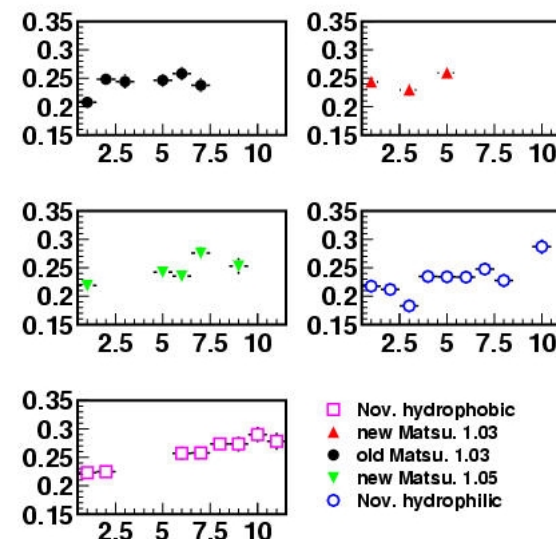
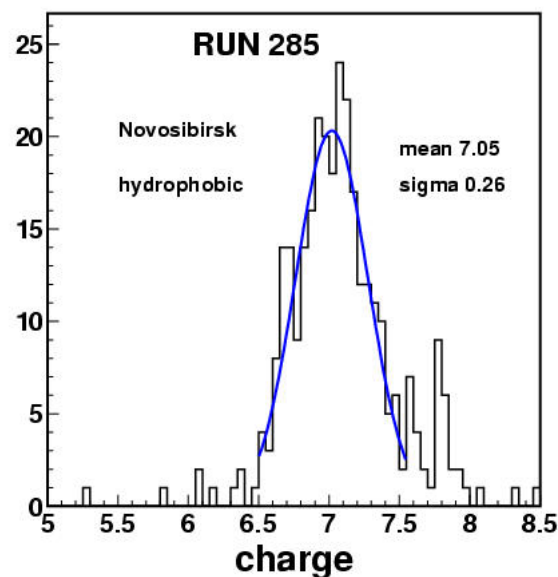
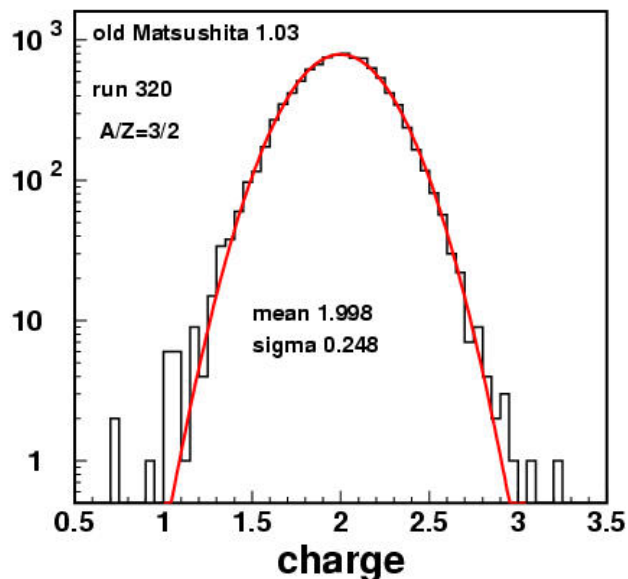
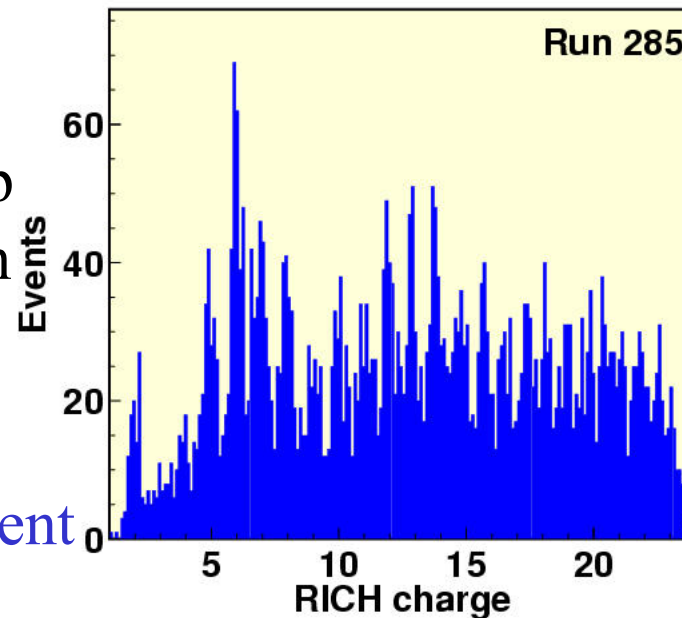


# PRELIMINARY RESULTS

## Z determination

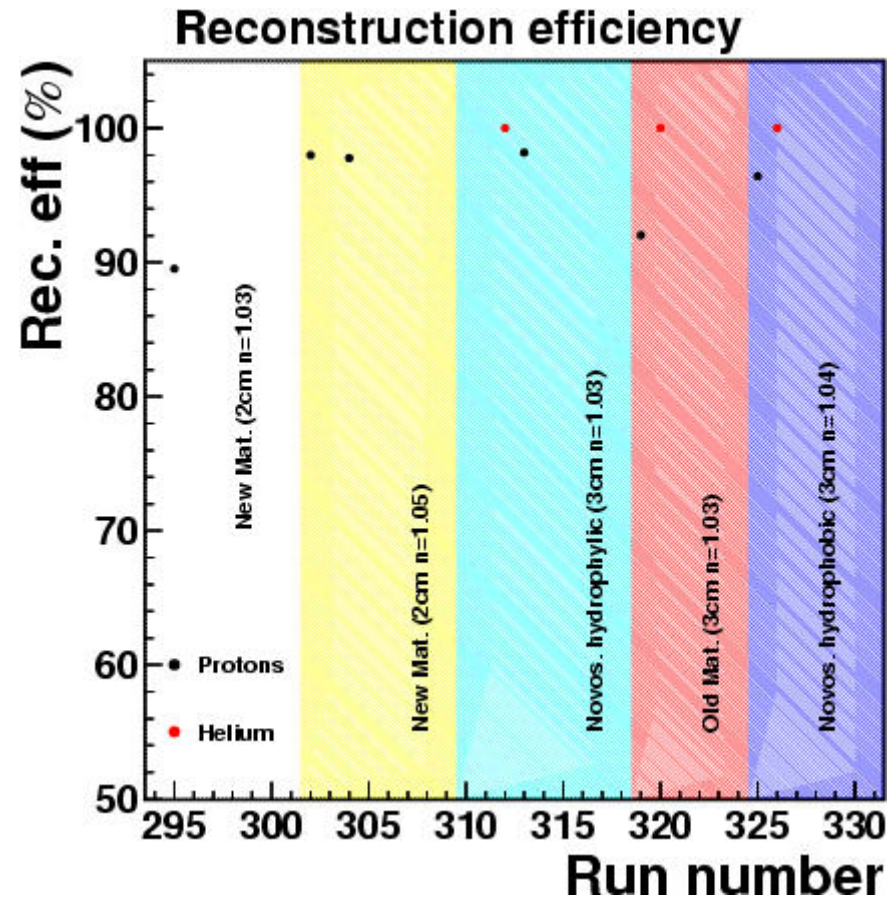
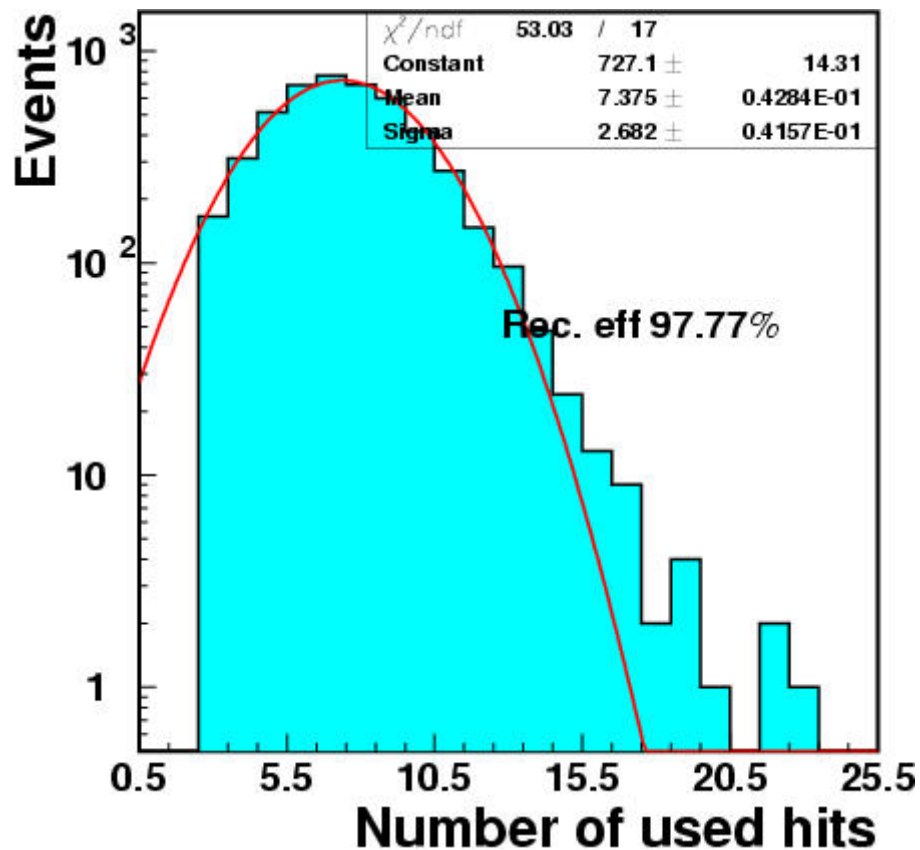
Although RICH charge peaks are visible up to  $Z \approx 20$ , scintillator charge determination limits a precise determination of RICH charge resolution & confusion for high  $Z$ .

RICH charge resolution for  $Z < 10$  is consistent with the expectations



# PRELIMINARY RESULTS

A first estimation of the RICH reconstruction efficiency can be derived from the extrapolation of the spectrum of used hits, which can be described to a good approximation by a Poisson distribution.

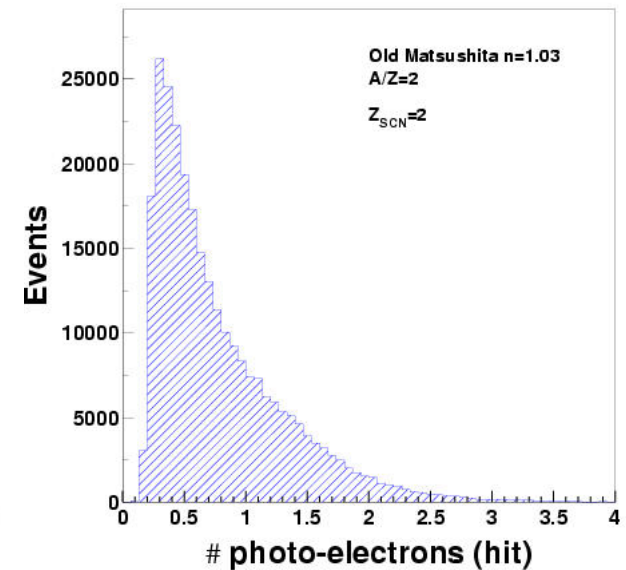
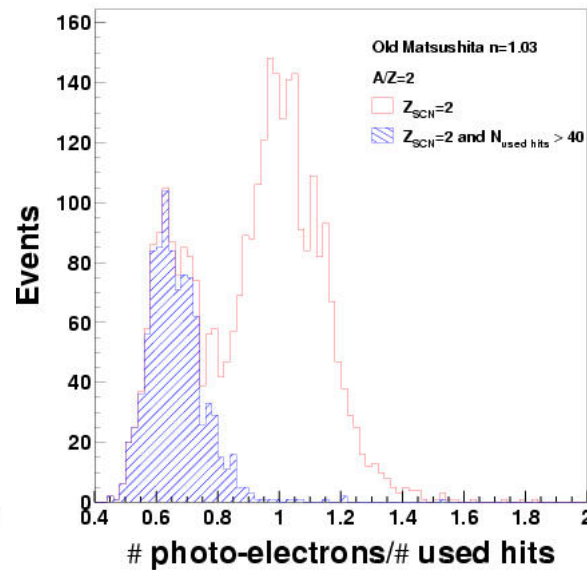
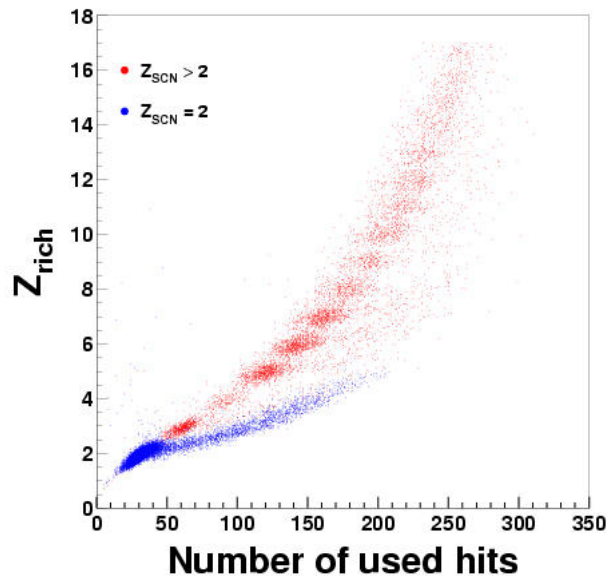


# STRANGE EVENTS

In  $Z_{\text{SCNT}}=2$  sample ( $A/Z = 2$  runs) 2 different types of RICH events:

- Standard Helium ( $N_{\text{HitsInRing}} < 40$ ,  $Q_{\text{HitInRing}} \approx 1$  p.e.)
- $N_{\text{HitsInRing}} > 40$ ,  $Q_{\text{HitInRing}} < 1$  p.e.)

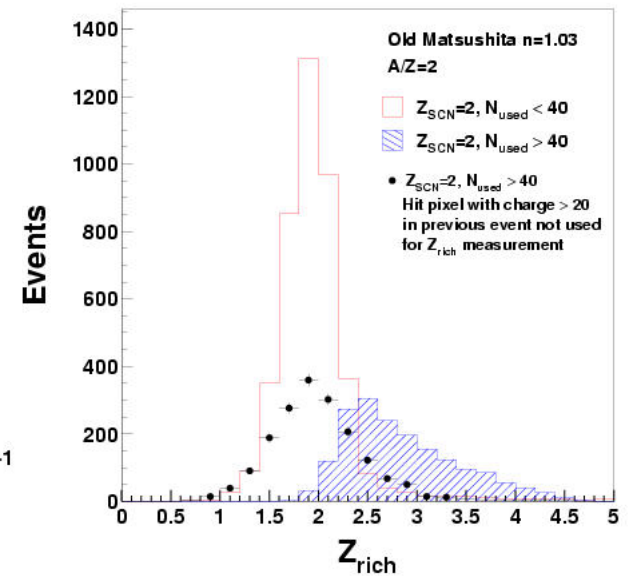
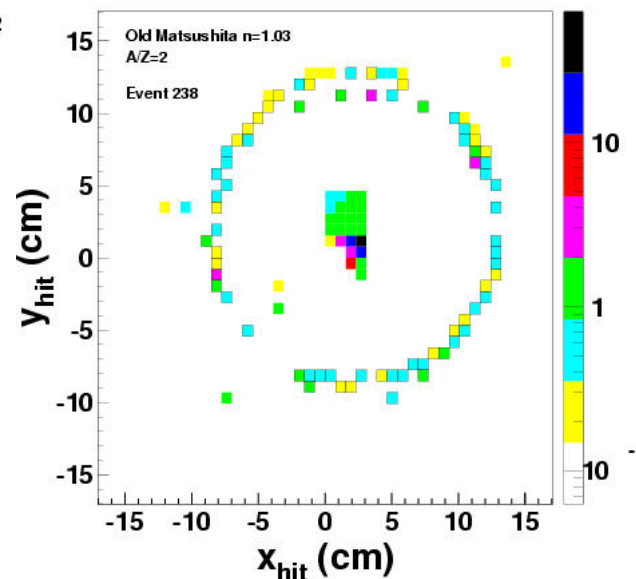
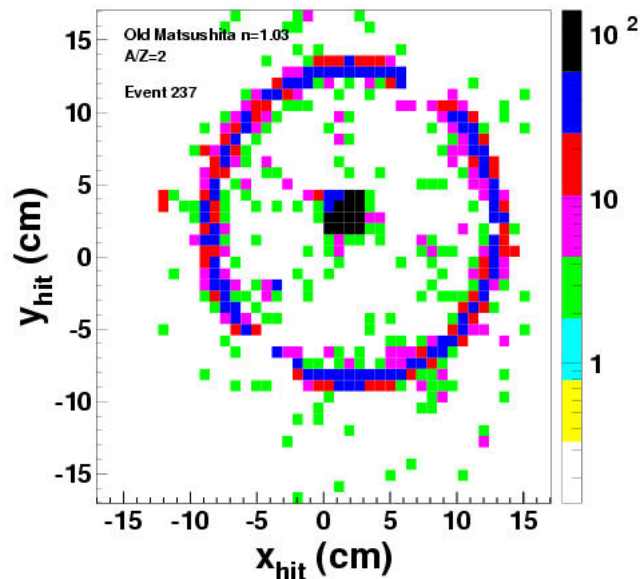
Actually, in sample B,  $Q_{\text{hit}}$  spectrum peak is below 1 p.e. !





# STRANGE EVENTS

- By comparing with previous trigger we see that the fired pixels in sample B are correlated with high charge hits in the previous event.
- After removing these coincident hits, sample B events show a nice  $Z_{\text{RICH}} = 2$  distribution.

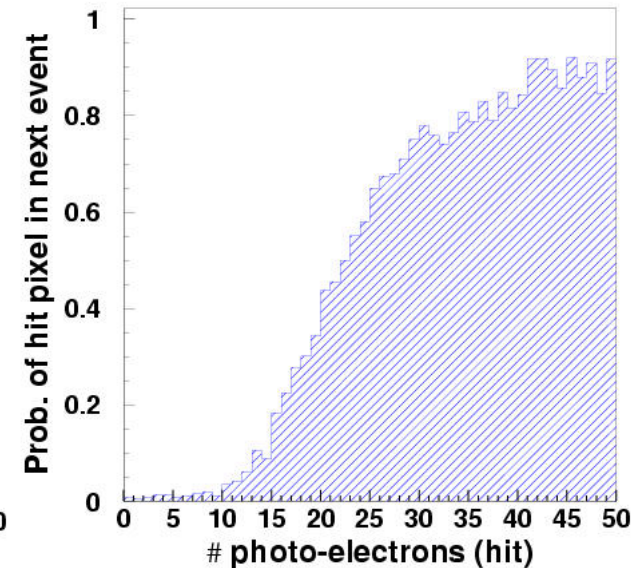
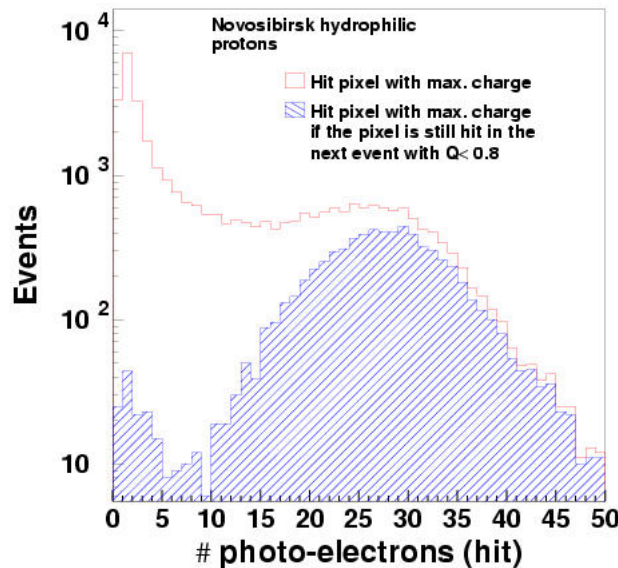
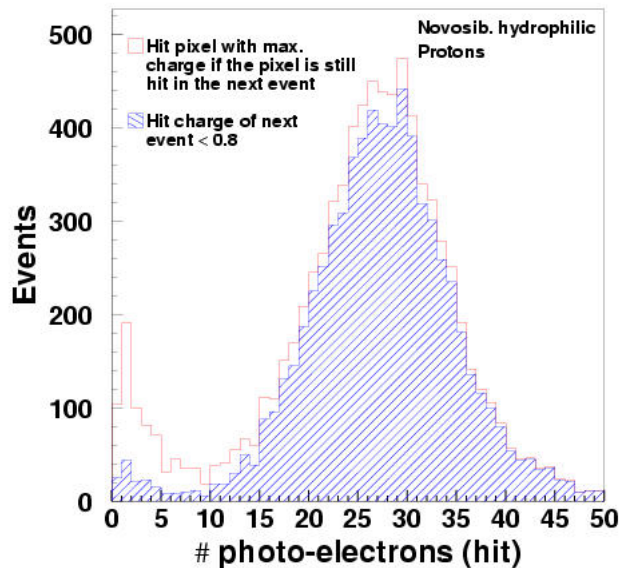
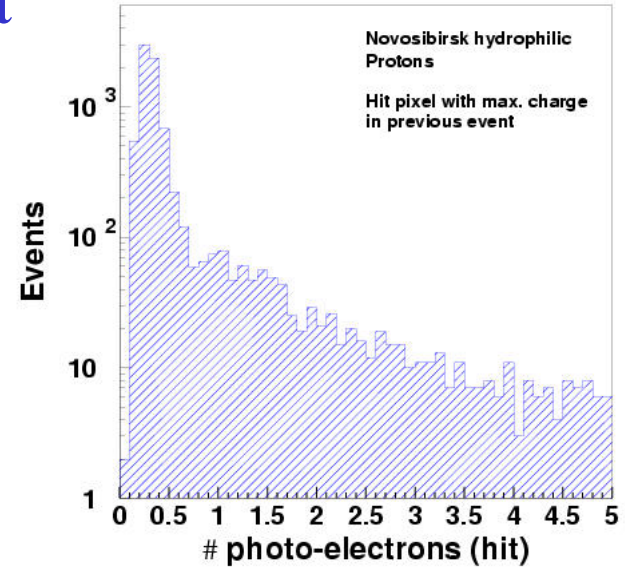


# STRANGE EVENTS

In the proton runs, we select in every event the pixel with the highest  $Q_{\text{hit}}$  ( $Q_{\text{hit}}^{\text{max}}$ ).

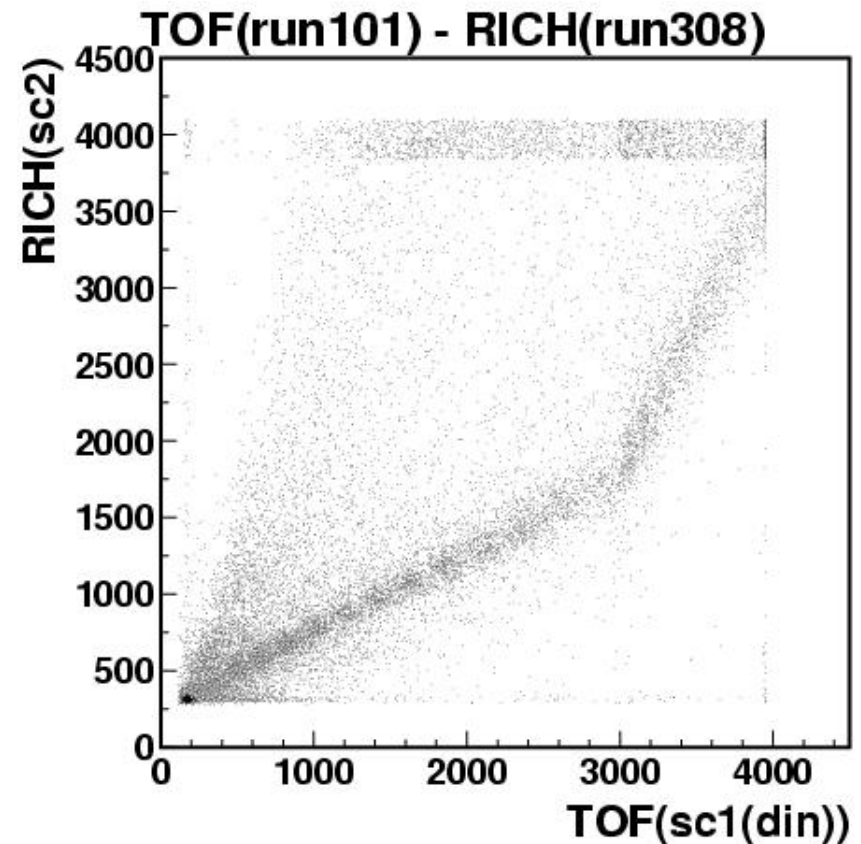
In the next event, the signal of that pixel, when fired, is strongly peaked at  $Q_{\text{hit}} < 1$ .

The probability of having a fired hit with low  $Q_{\text{hit}}$  increases with increasing  $Q_{\text{hit}}^{\text{max}}$ .



# NEXT STEPS

- **Calibration verification (WC, SCNT...)**
- **Ntuple availability**
- **MC production**
- **Analysis extension to the full data sample including all the radiators**



Additional background rejection and identification capabilities can be obtained by correlating RICH data with Tracker and TOF samples using the common event number.

# CONCLUSIONS

- First look at test beam data agrees with the system expected performances regarding velocity resolution and dynamic range.
- Detailed comparison with simulation is needed to perform the final radiator choice.
- Correlation with TOF and Tracker data may be needed to reject backgrounds and to improve signal selection.
- Two unexpected effects have been detected:
  - Inconsistent LED/Offline PMT Gain calibration
  - “Persistent signals” in consecutive events